



The astrophinformatics Laboratory at CMM



Jaime San Martín
Professor
Ph.D.
Mathematics



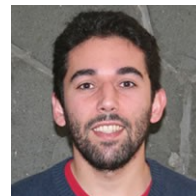
Eduardo Vera
Scientist
Ph.D. Physics



Guillermo Cabrera
Engineer
Ph.D. Comp. Sc.



Pablo Román
Postdoc
Ph.D. Systems
Eng.



Francisco Förster
Scientist
Ph.D. Astronomy



J. Carlos Maureira
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Santiago González-Gaitán
Postdoc
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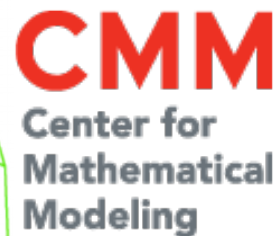
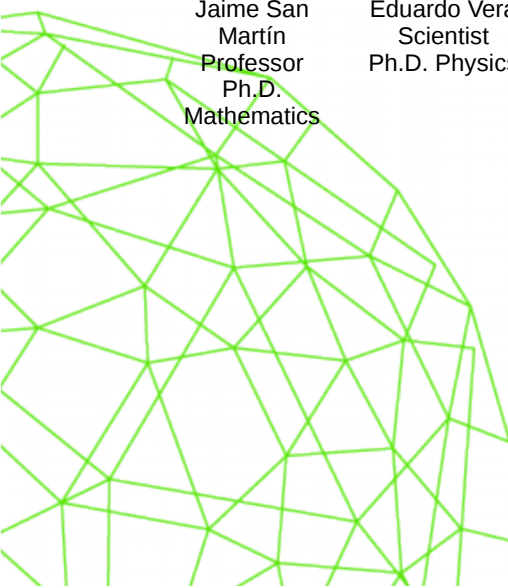


Pablo Huijse
Postdoc
Ph.D. Elec. Eng.



Pablo Estévez
Professor
Ph.D. Elec. Eng.

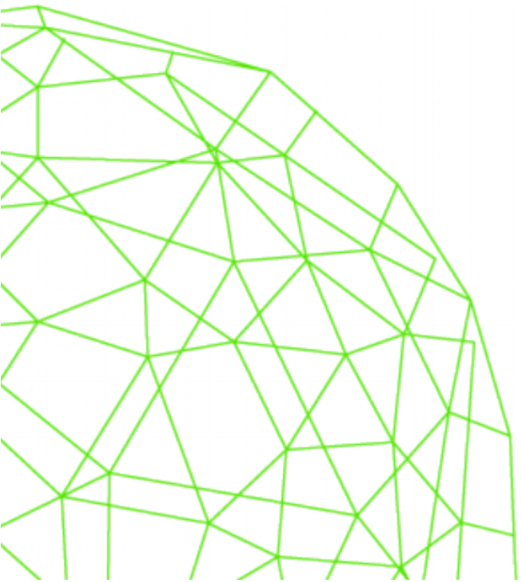
and many more collaborators...



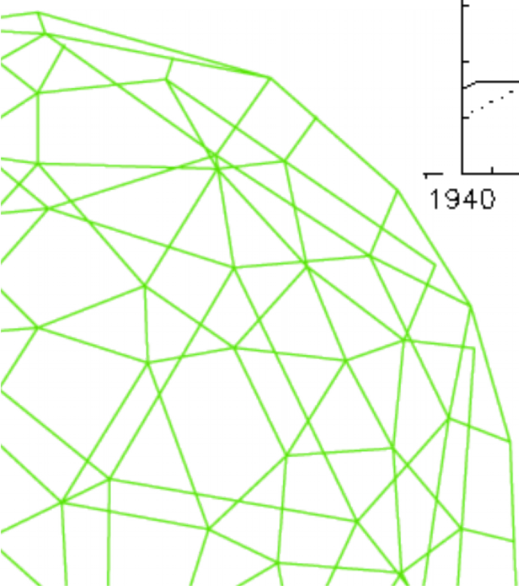
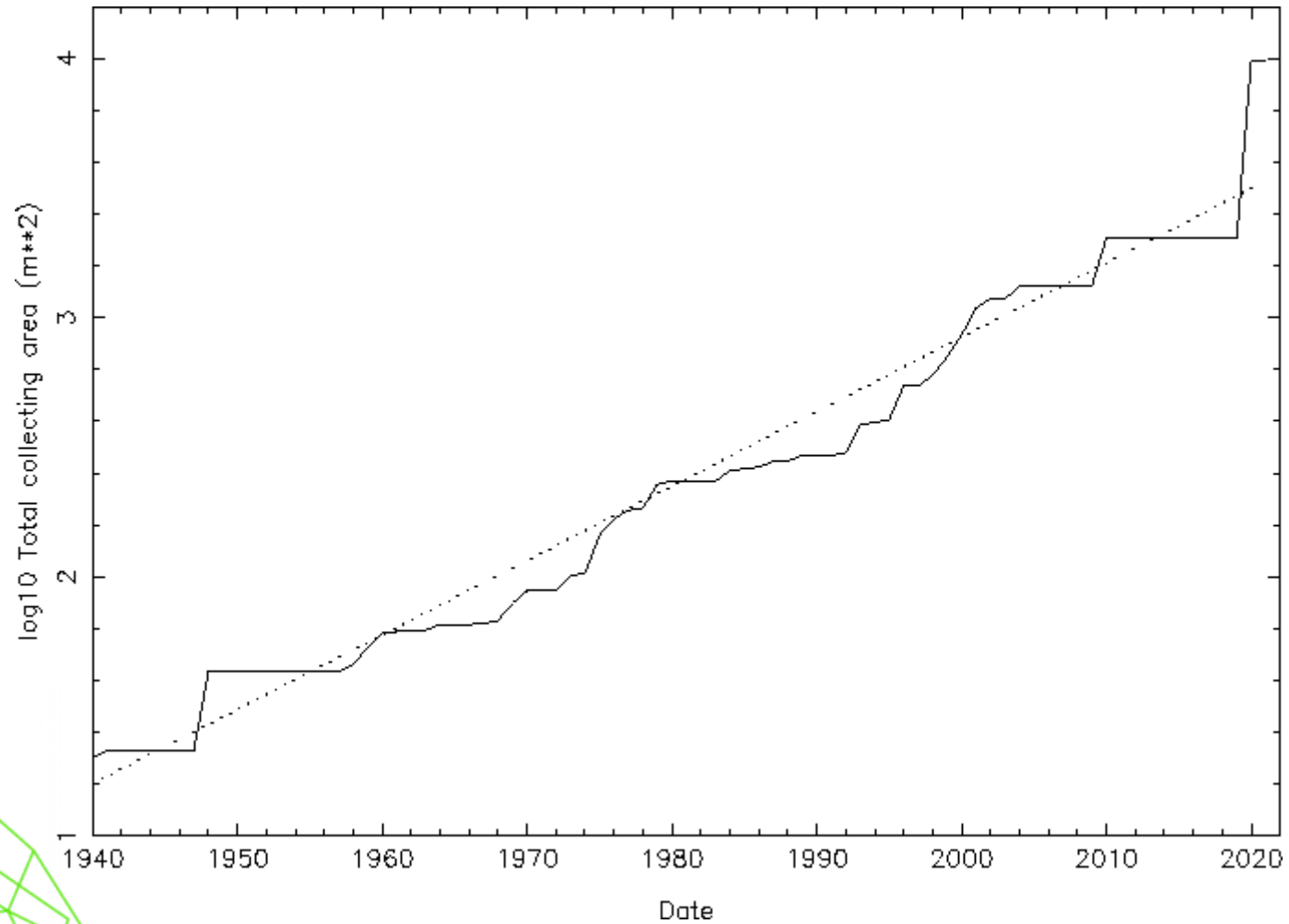
Astroinformatics Laboratory

- Focus in astronomical problems that require the analysis of **large volumes of data**.
- Efficient algorithms, statistics, data handling methods, visualization, inverse problems, image processing, machine learning, high performance computing.
- **Interdisciplinary** focus gives **flexibility** to move into new developing areas of astronomy.
- Focus on **data-driven** and well defined problems.
- e.g. real-time, high-cadence survey for transient objects (**DECam**); automatic classification of galaxies (**LSST**), characterization of light curves (**DECam**, **LSST**) and image synthesis in interferometric observations (**ALMA**).

Astronomy in the next decade: Big Data & Big Telescopes



Big Telescopes



Great Paris Exhibition Telescope
(lens at the same scale)
Paris, France (1900)

Yerkes Observatory
(40" refractor lens at the same scale)
Williams Bay, Wisconsin (1893)

Hooker (100")
Mt Wilson, California (1917)

Hale (200")
Mt Palomar, California (1948)

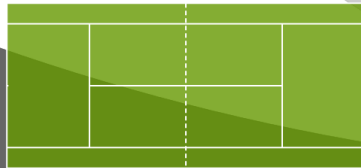
Multi Mirror Telescope
(1979-1998)
Mount Hopkins, Arizona

BTA-6 (Large Altazimuth Telescope)
Zelenchuksky, Russia (1975)

Large Zenith Telescope
British Columbia, Canada (2003)

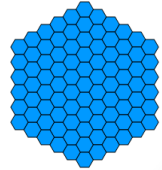
Gaia
Earth-Sun L2 point (2014)

James Webb Space Telescope
Earth-Sun L2 point (planned 2018)

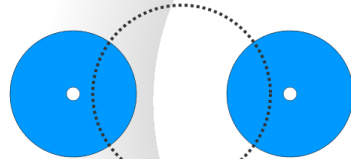


Tennis court at the same scale

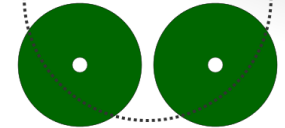
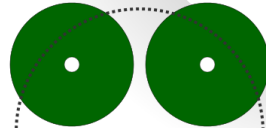
Large Sky Area Multi-Object Fiber Spectroscopic Telescope
Hebei, China (2009)



Hobby-Eberly Telescope
Davis Mountains, Texas (1996)



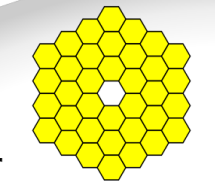
Large Binocular Telescope
Mount Graham, Arizona (2005)



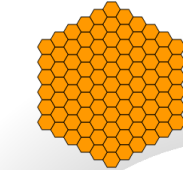
Very Large Telescope
Cerro Paranal, Chile (1998-2000)



Magellan Telescopes
Las Campanas, Chile (2000/2002)



Gran Telescopio Canarias
La Palma, Canary Islands, Spain (2007)



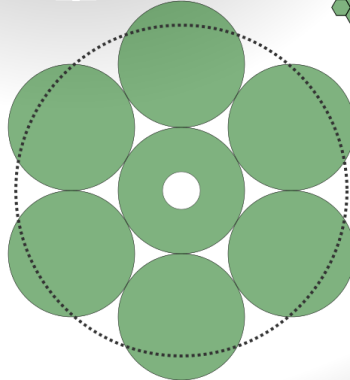
Southern African Large Telescope
Sutherland, South Africa (2005)



Gemini South
Cerro Pachón, Chile (2000)



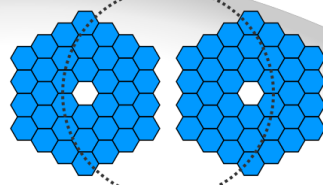
Large Synoptic Survey Telescope
El Peñón, Chile (planned 2020)



Giant Magellan Telescope
Las Campanas Observatory, Chile (planned 2020)

Overwhelmingly Large Telescope
(cancelled)

Arecibo radio telescope at the same scale



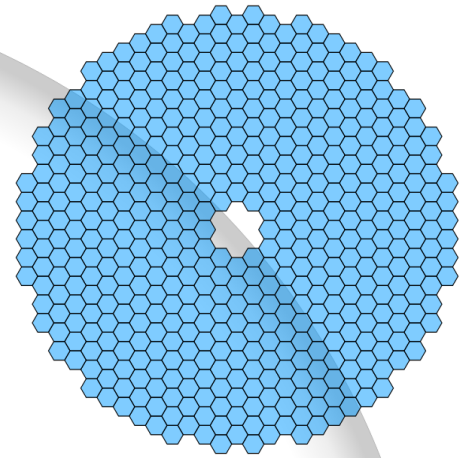
Keck Telescope
Mauna Kea, Hawaii (1993/1996)



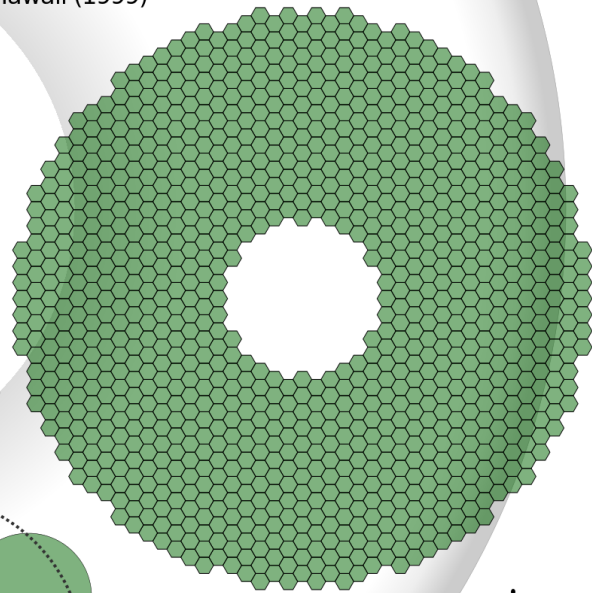
Gemini North
Mauna Kea, Hawaii (1999)



Subaru Telescope
Mauna Kea, Hawaii (1999)

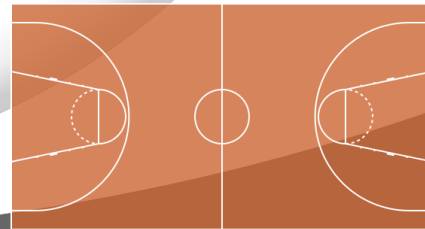
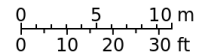


Thirty Meter Telescope
Mauna Kea, Hawaii (planned 2022)

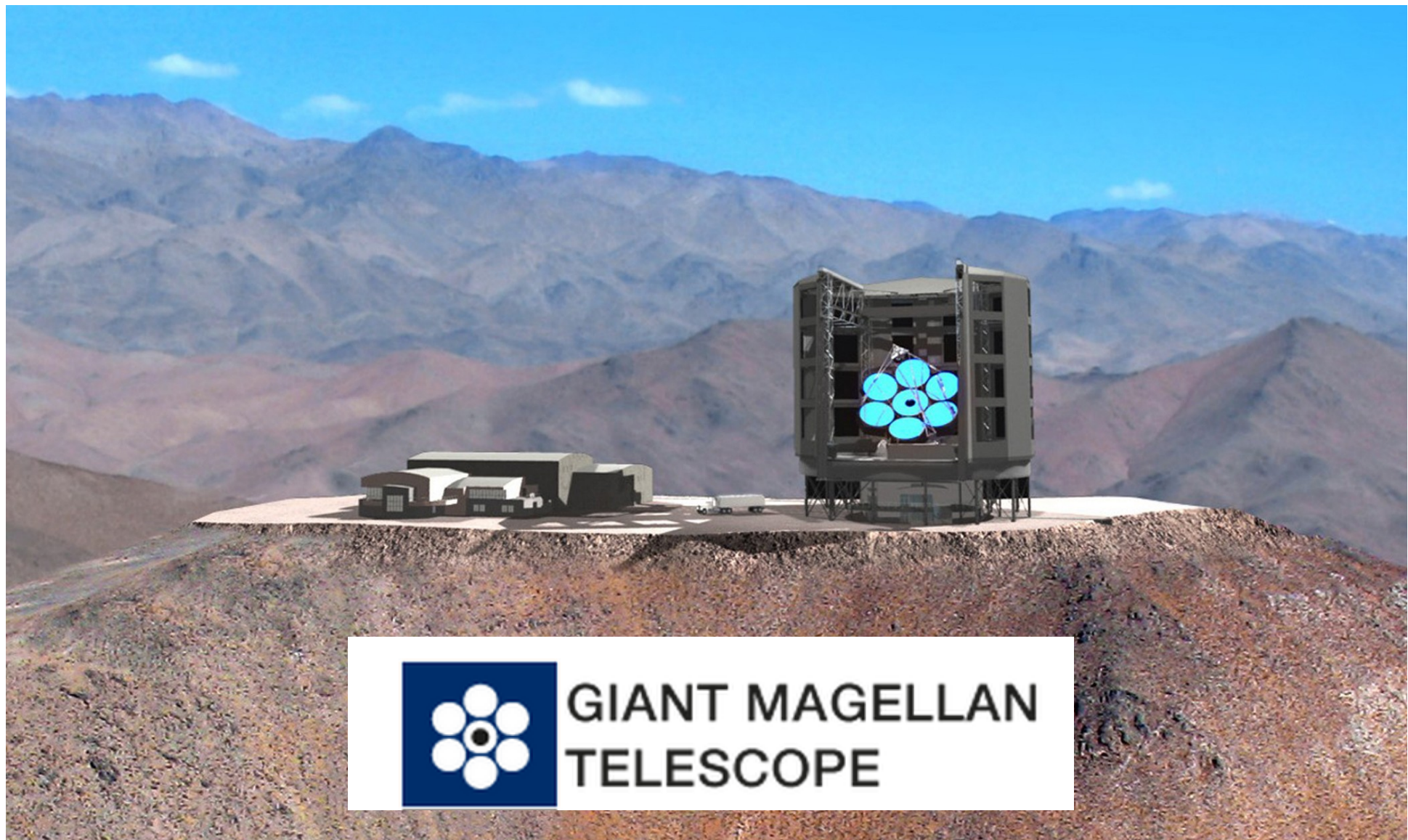


European Extremely Large Telescope
Cerro Armazones, Chile (planned 2022)

Human at the same scale

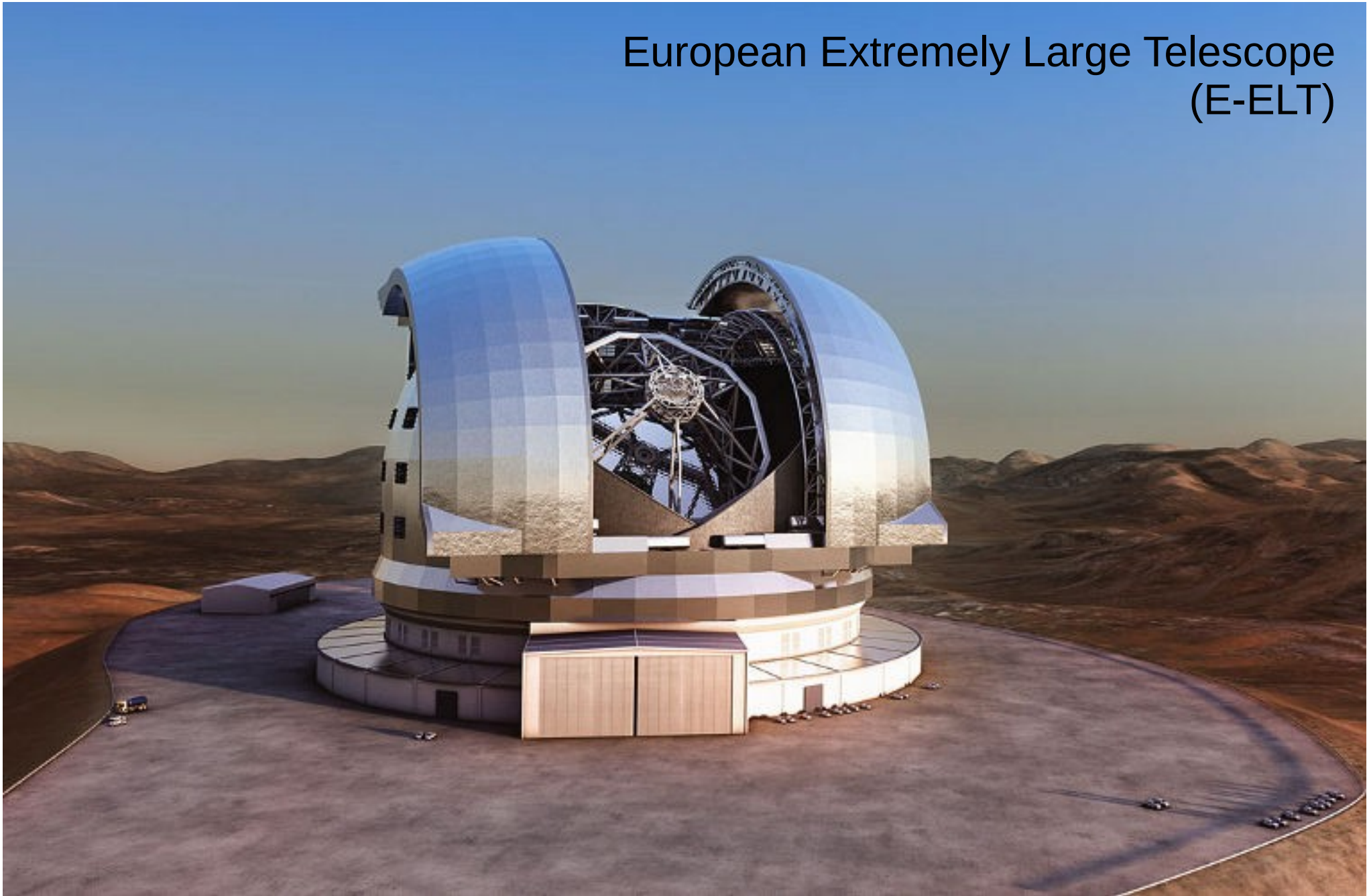


Basketball court at the same scale

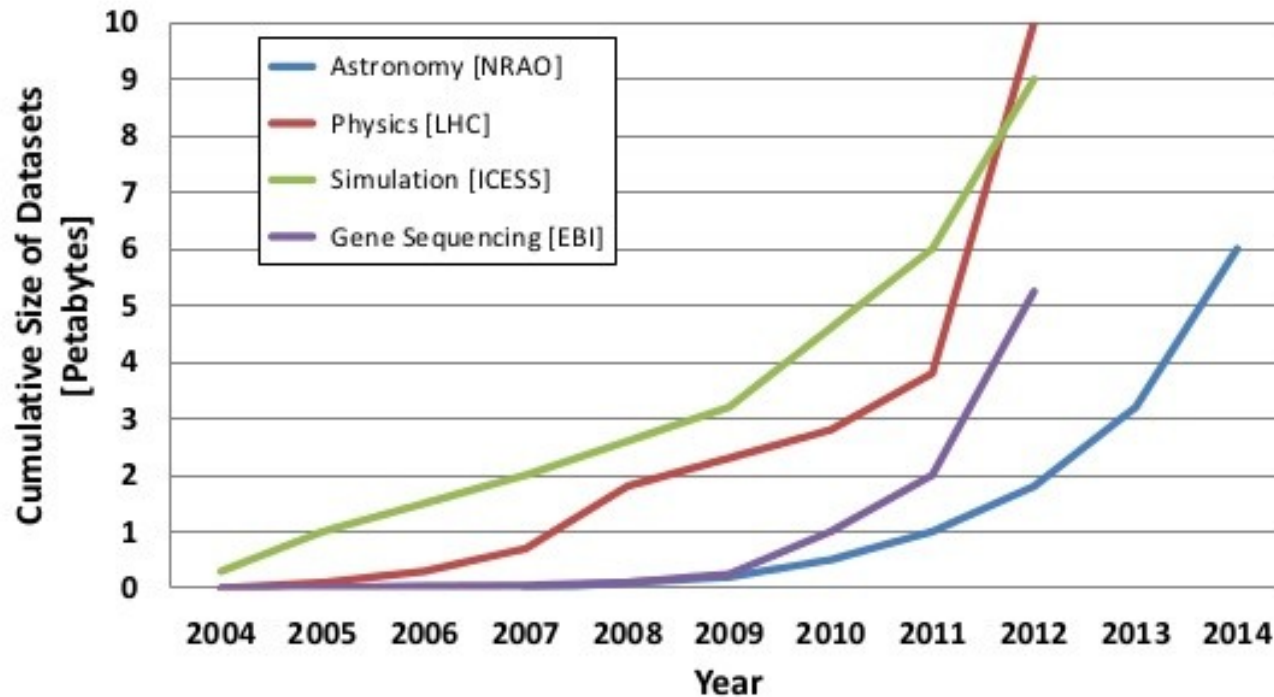


GIANT MAGELLAN
TELESCOPE

European Extremely Large Telescope (E-ELT)



Big Data in astronomy



Scientific Data Grows Exponentially!

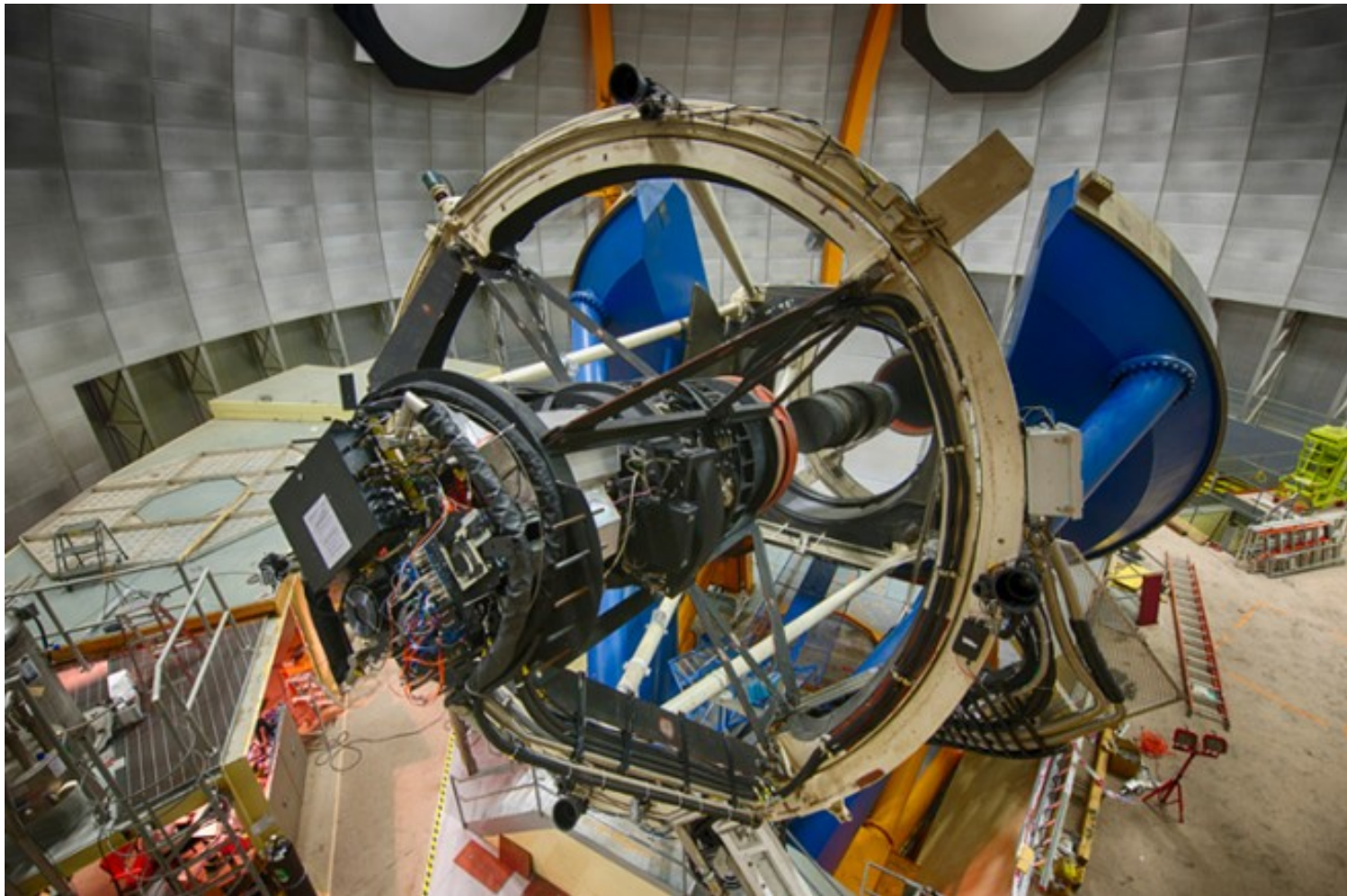
3



~0.2 PB/yr



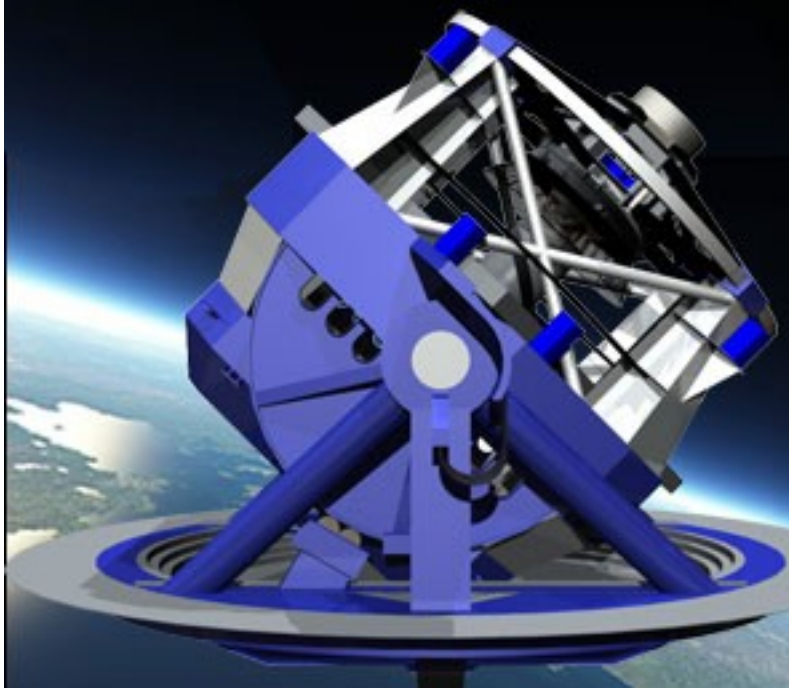
Dark Energy Camera (DECam)



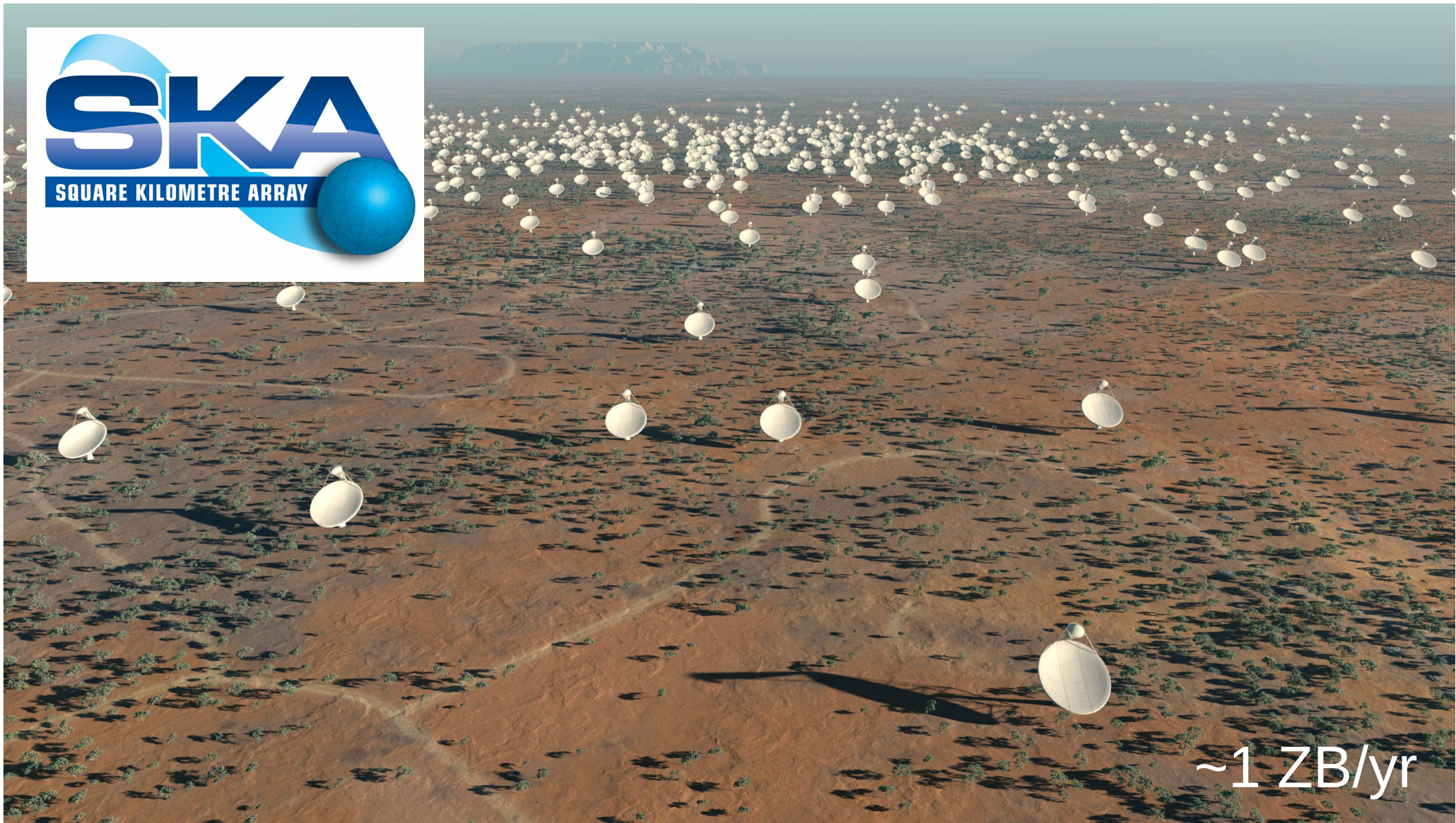
~0.4 PB/yr

LST

Large Synoptic Survey Telescope



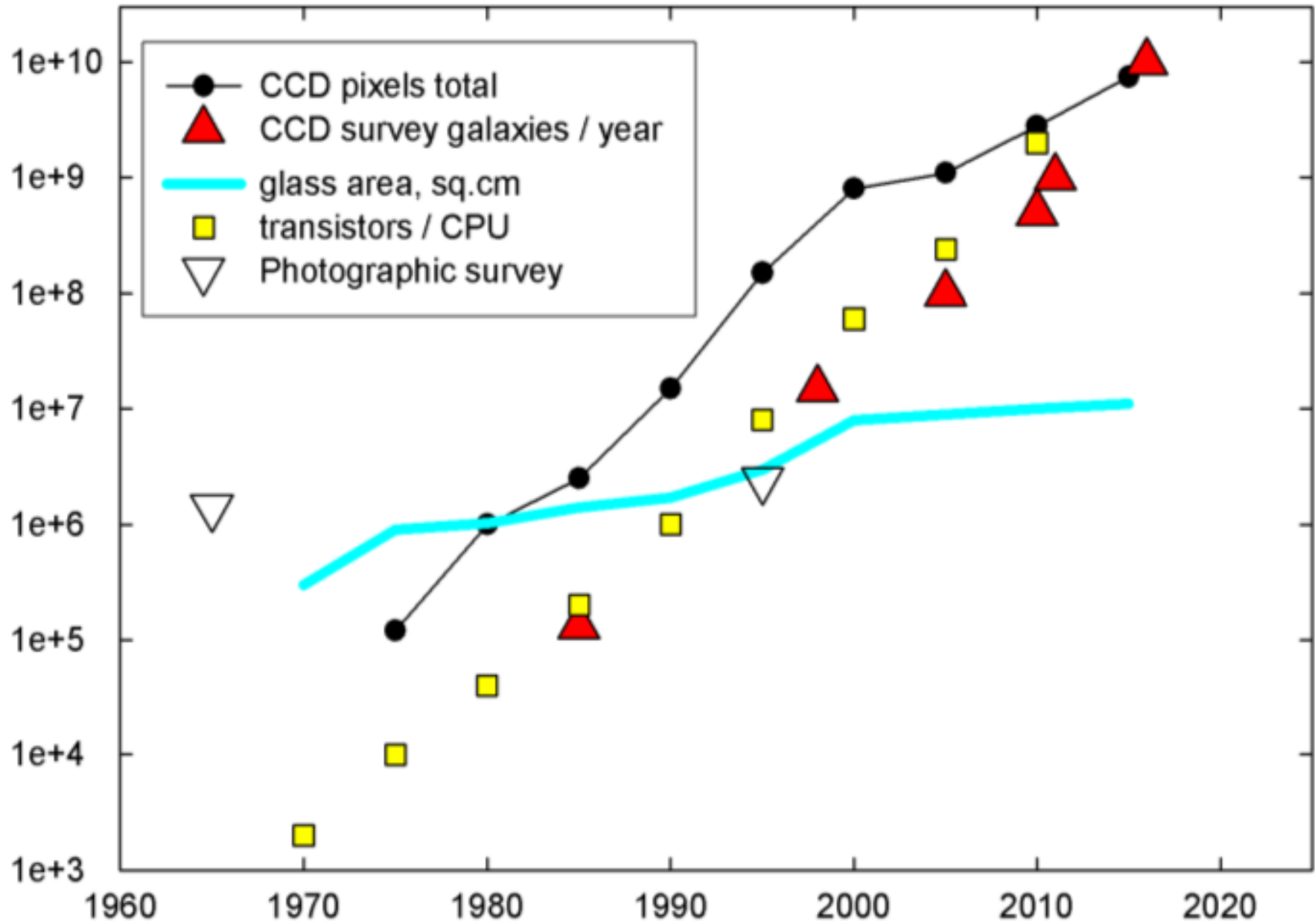
~5 PB/yr



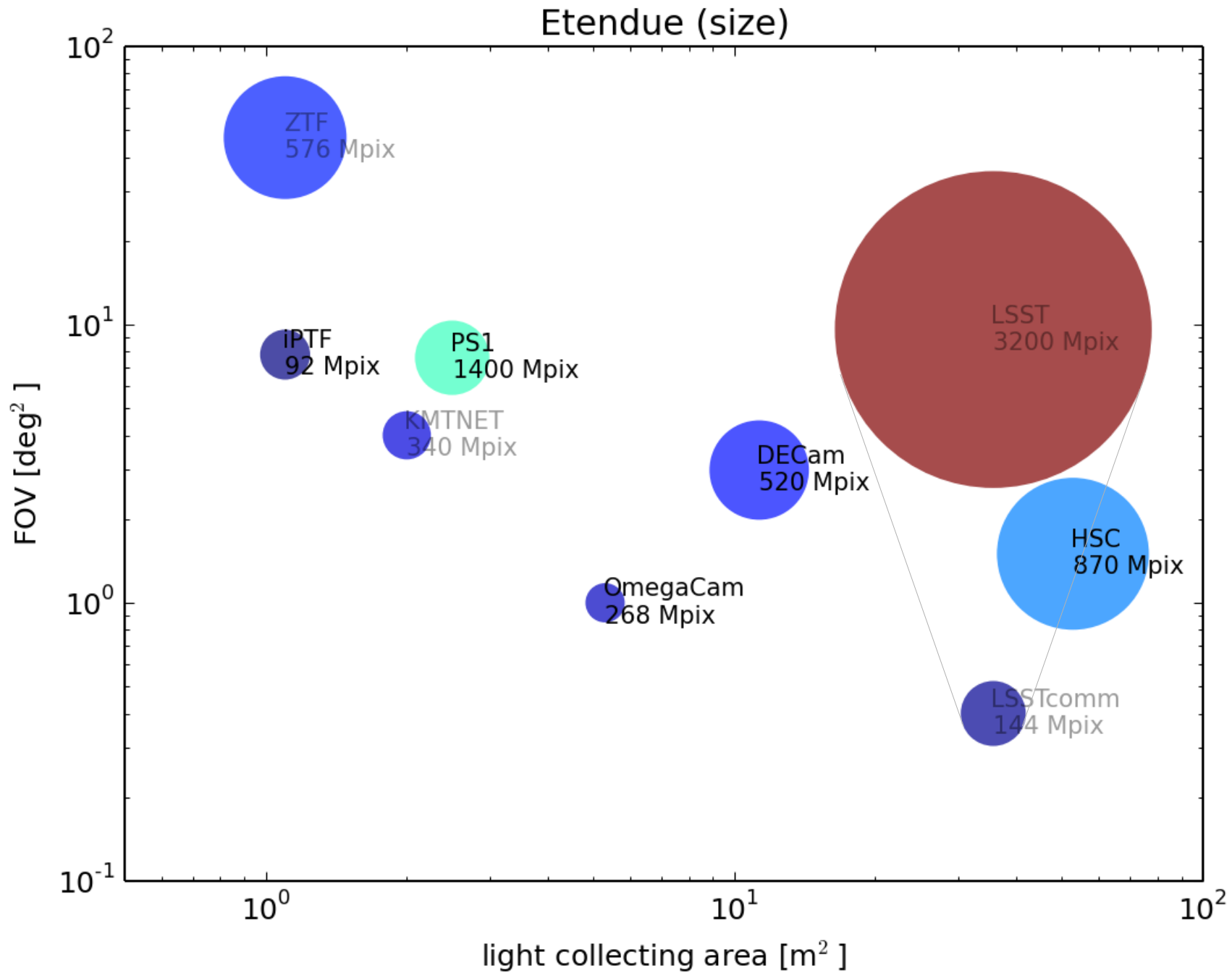
1 Zettabyte = 10^3 ExaBytes = 10^6 PB



Data volume > telescope size



Etendue and number of pixels



The large Synoptic Survey Telescope (LSST)

Time domain science

- Nova, supernova, GRBs
- Source characterization
- Instantaneous discovery

Census of the Solar System

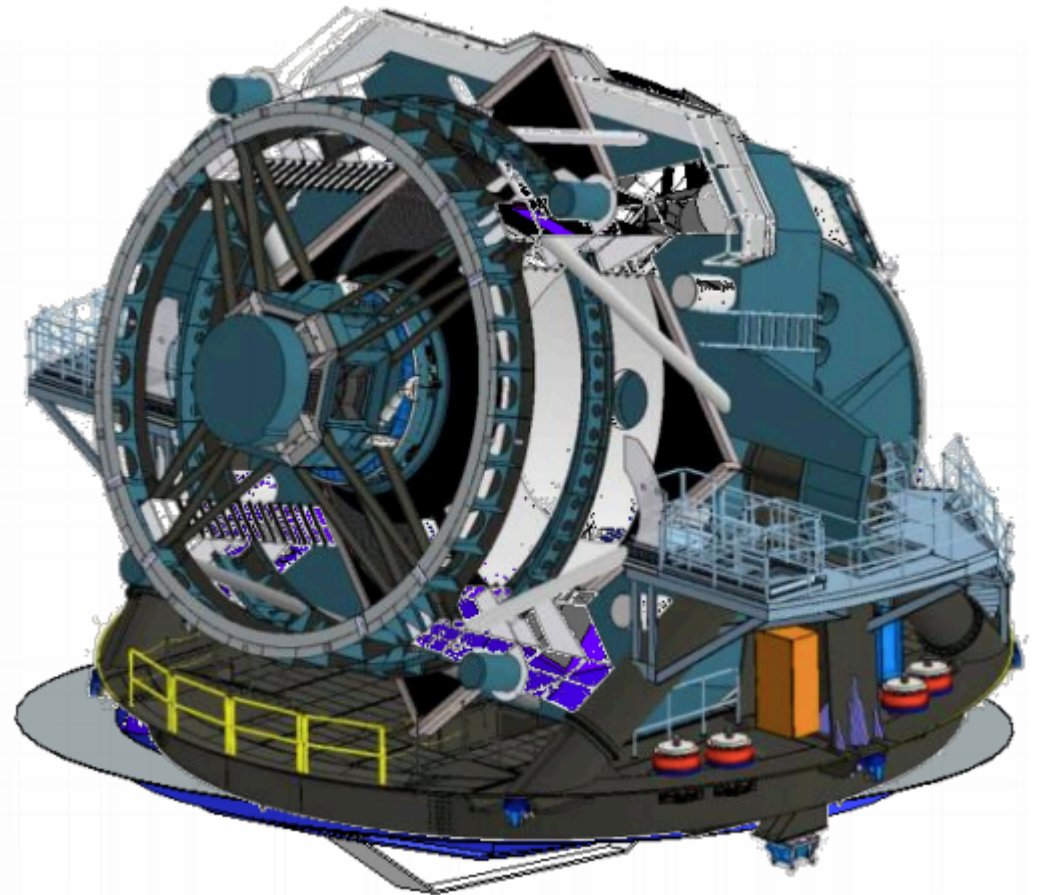
- NEOs, MBAs, Comets
- KBOs, Oort Cloud

Mapping the Milky Way

- Tidal streams
- Galactic structure

Dark energy and dark matter

- Strong Lensing
- Weak Lensing
- Constraining the nature of dark energy



The large Synoptic Survey Telescope (LSST)

Traditional scientists

What data should I collect to test an hypothesis?

Data-driven scientist

What theories can I test given the data I already have?

Success in research will depend on the ability to mine knowledge from data.



The large Synoptic Survey Telescope (LSST)

LEVEL 1: Nightly alerts

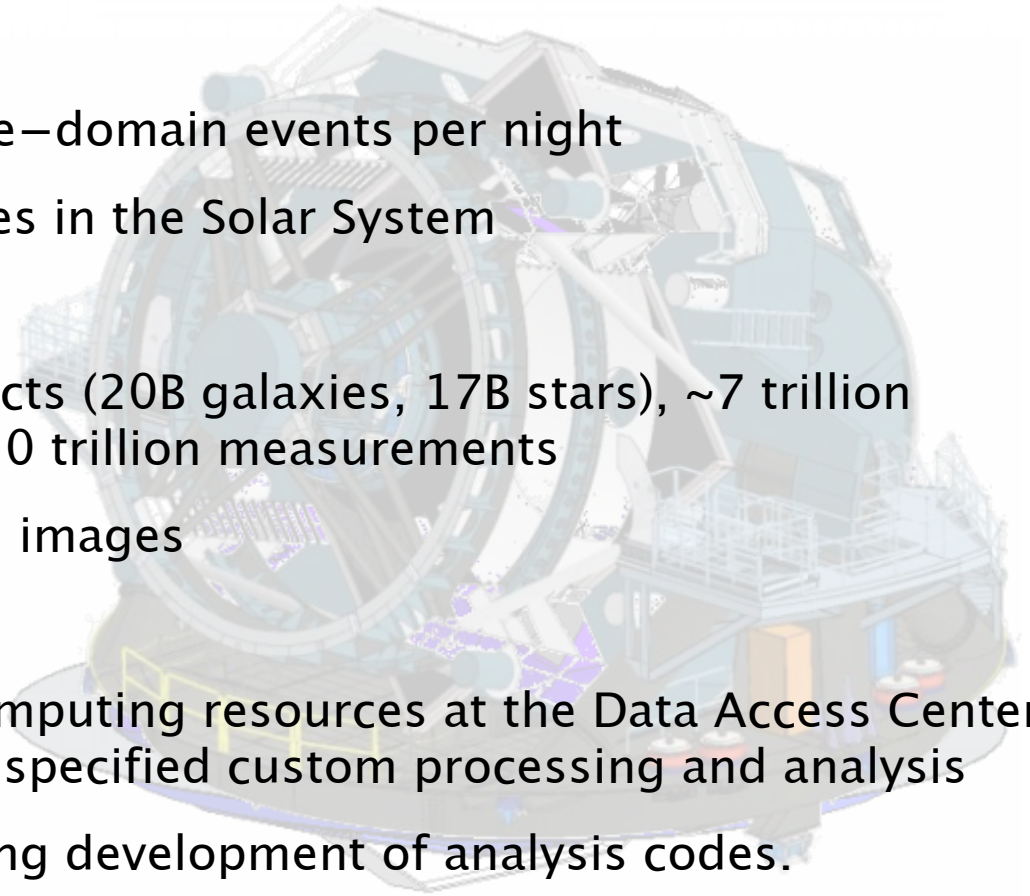
~10 million time–domain events per night
~6 million bodies in the Solar System

LEVEL 2: Yearly data releases

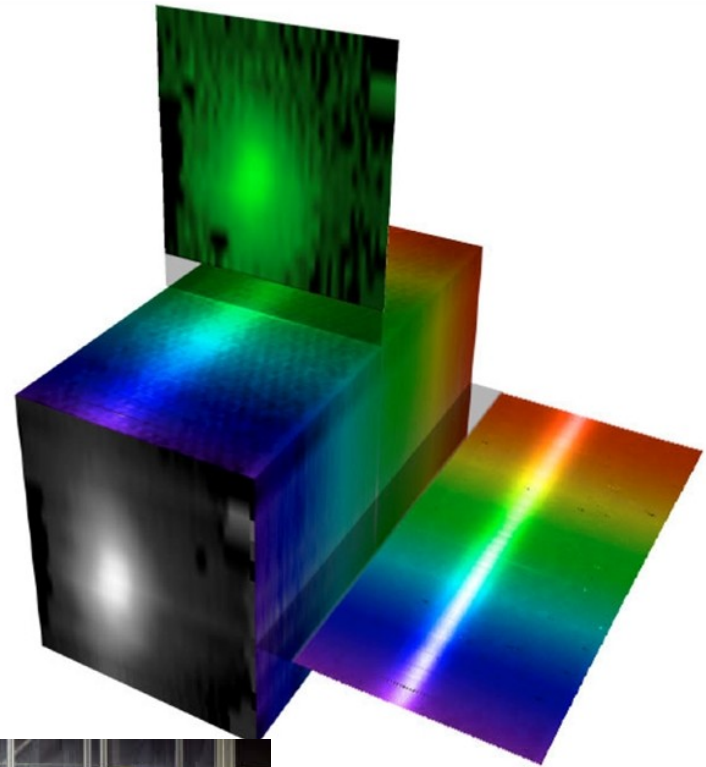
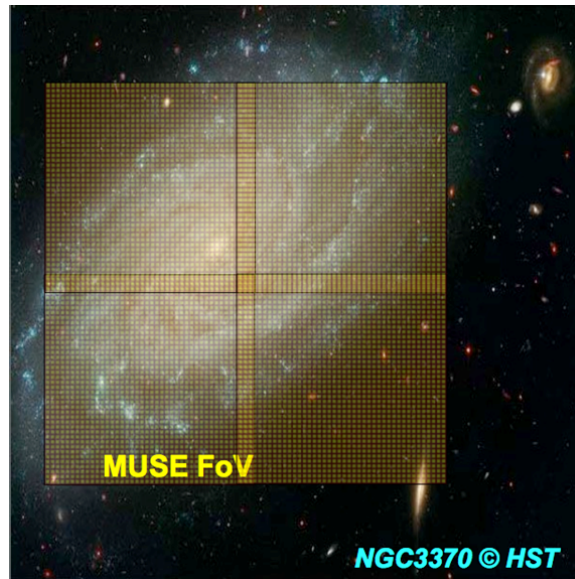
~37 billion objects (20B galaxies, 17B stars), ~7 trillion sources, and ~30 trillion measurements
Deep co–added images

LEVEL 3: Community projects

Services and computing resources at the Data Access Centers to enable user–specified custom processing and analysis
Software enabling development of analysis codes.



Integral Field Units, e.g. Multi-Unit Spectroscopic Explorer (MUSE)



Internet submarine cables



Veracity

Curation

Search

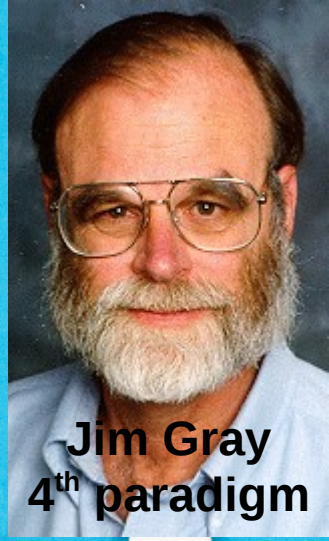
Capture

Storage

Velocity

BIG DATA

Sharing



Jim Gray
4th paradigm

Variety

Analysis

Volume

Visualization

Transfer

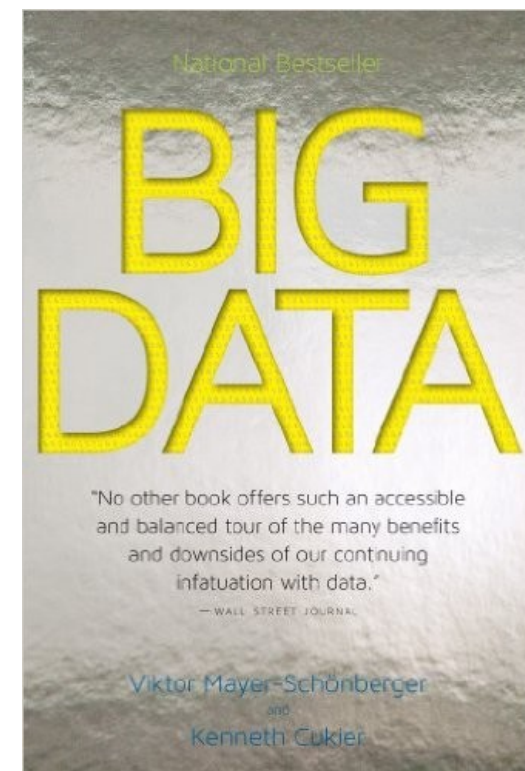
Big data

“Big Data is all about seeing and understanding the relations within and among pieces of information...”

“It is about three major shifts of mindset ... the ability to analyze **vast amounts of data** about a topic ... willingness to embrace **data's real-world messiness** ... growing respect for **correlations** rather than a continuous quest for elusive **causality**.”

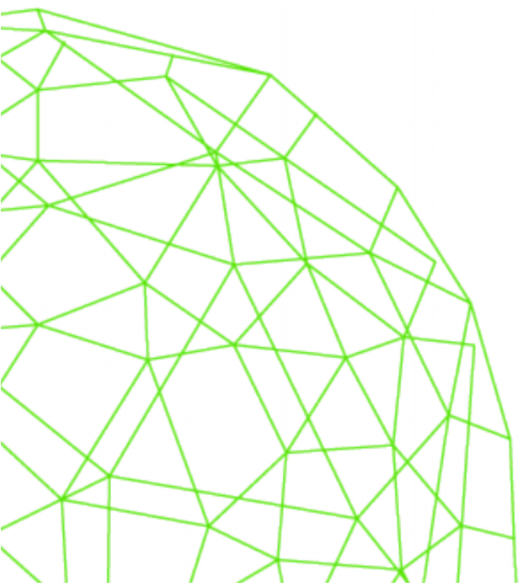
“Simple models and a lot of data trump over more elaborate models on less data”

Peter Norvig, Google artificial intelligence.

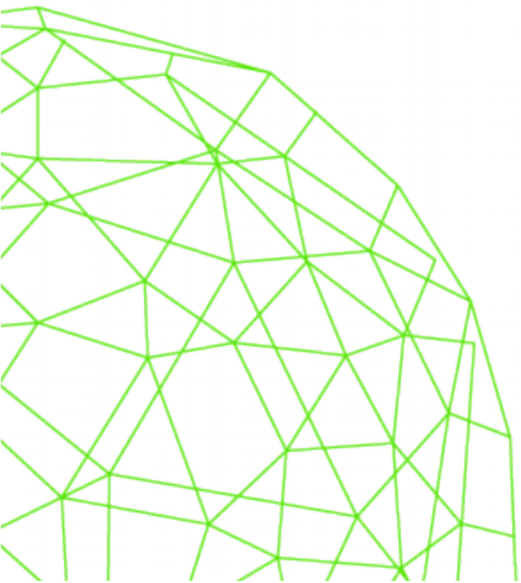
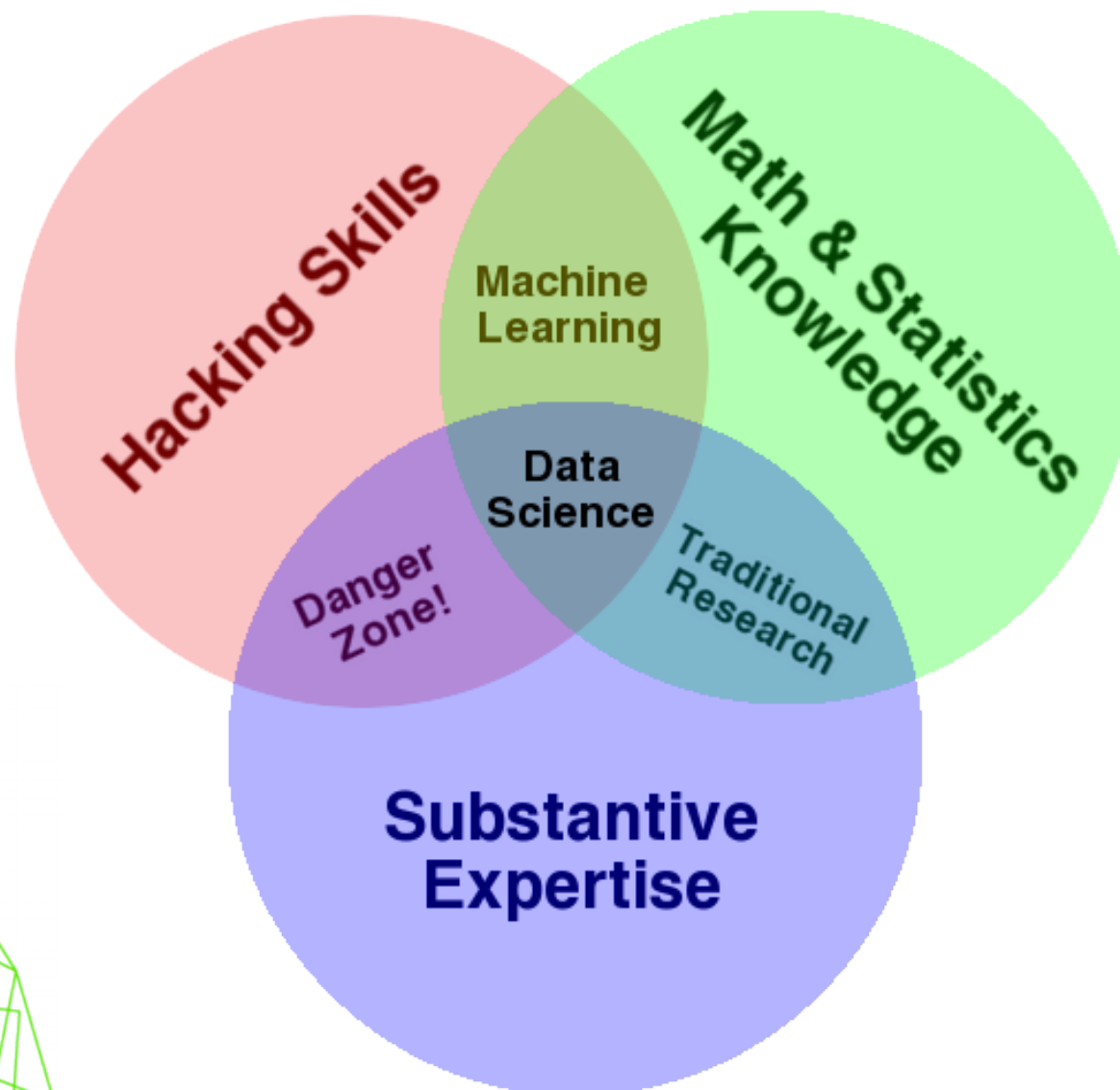


"Big data is like teenage sex: everyone talks about it, nobody really knows how to do it, everyone thinks everyone else is doing it, so everyone claims they are doing it..."

Dan Ariely,
Duke University



Data science



**Harvard
Business
Review**

2012

DATA

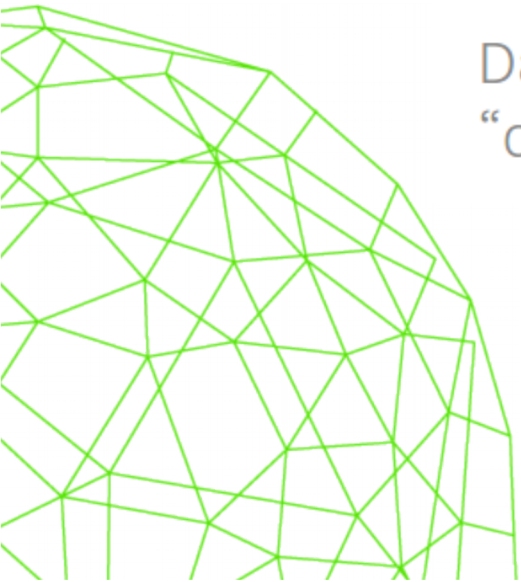
Data Scientist: The Sexiest Job of the 21st Century

by **Thomas H. Davenport** and **D.J. Patil**

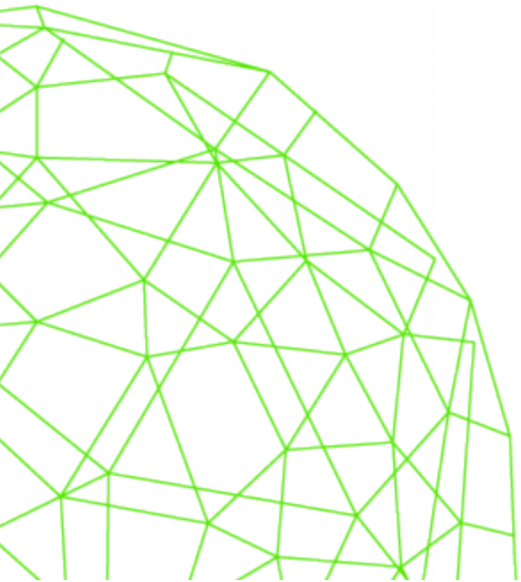
FROM THE OCTOBER 2012 ISSUE

The shortage of data scientists is becoming a serious constraint in some sectors.

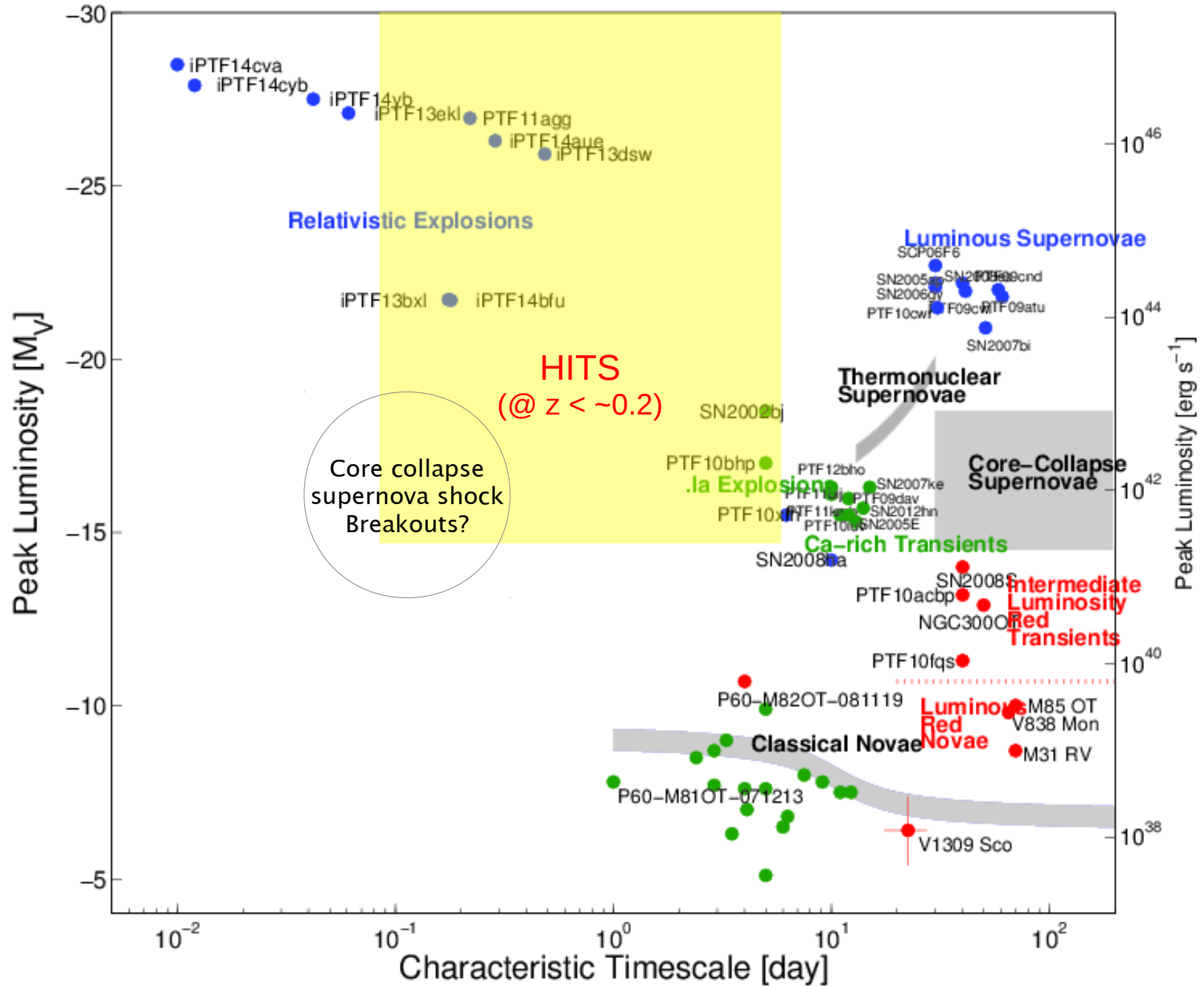
Data scientists today are akin to the Wall Street “quants” of the 1980s and 1990s.



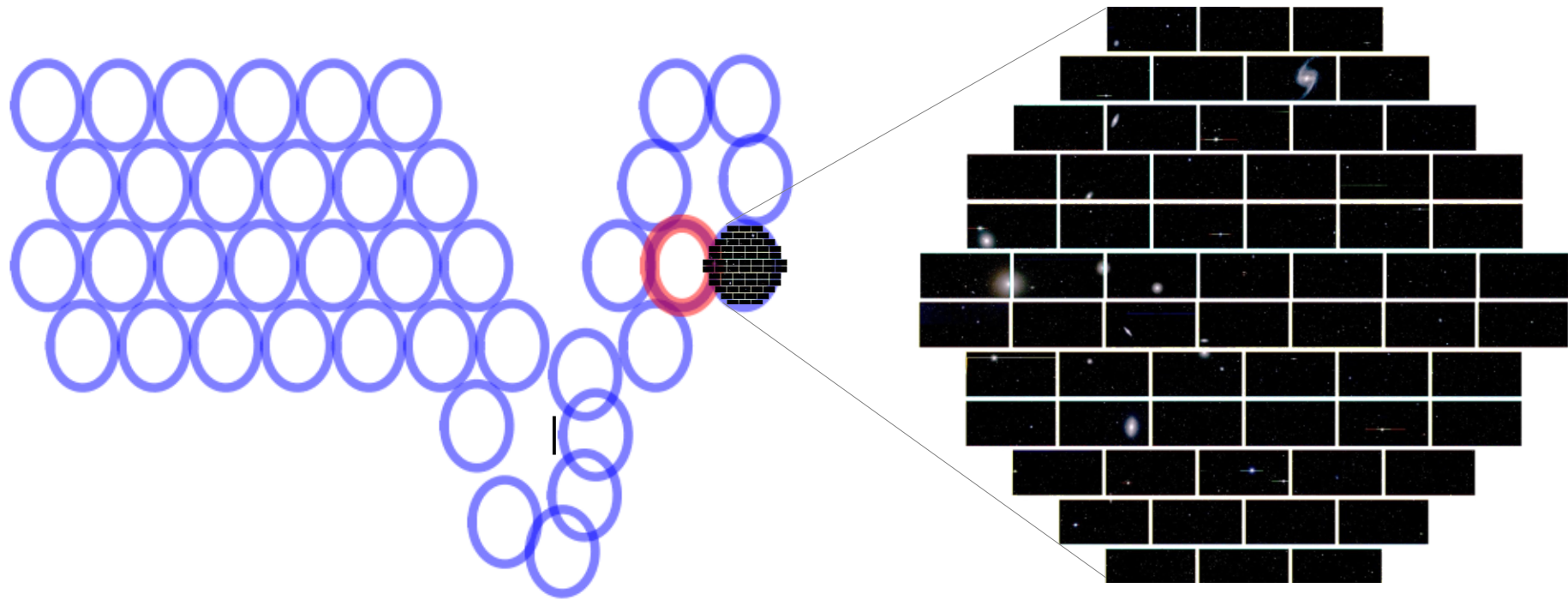
Astroinformatics Laboratory Projects

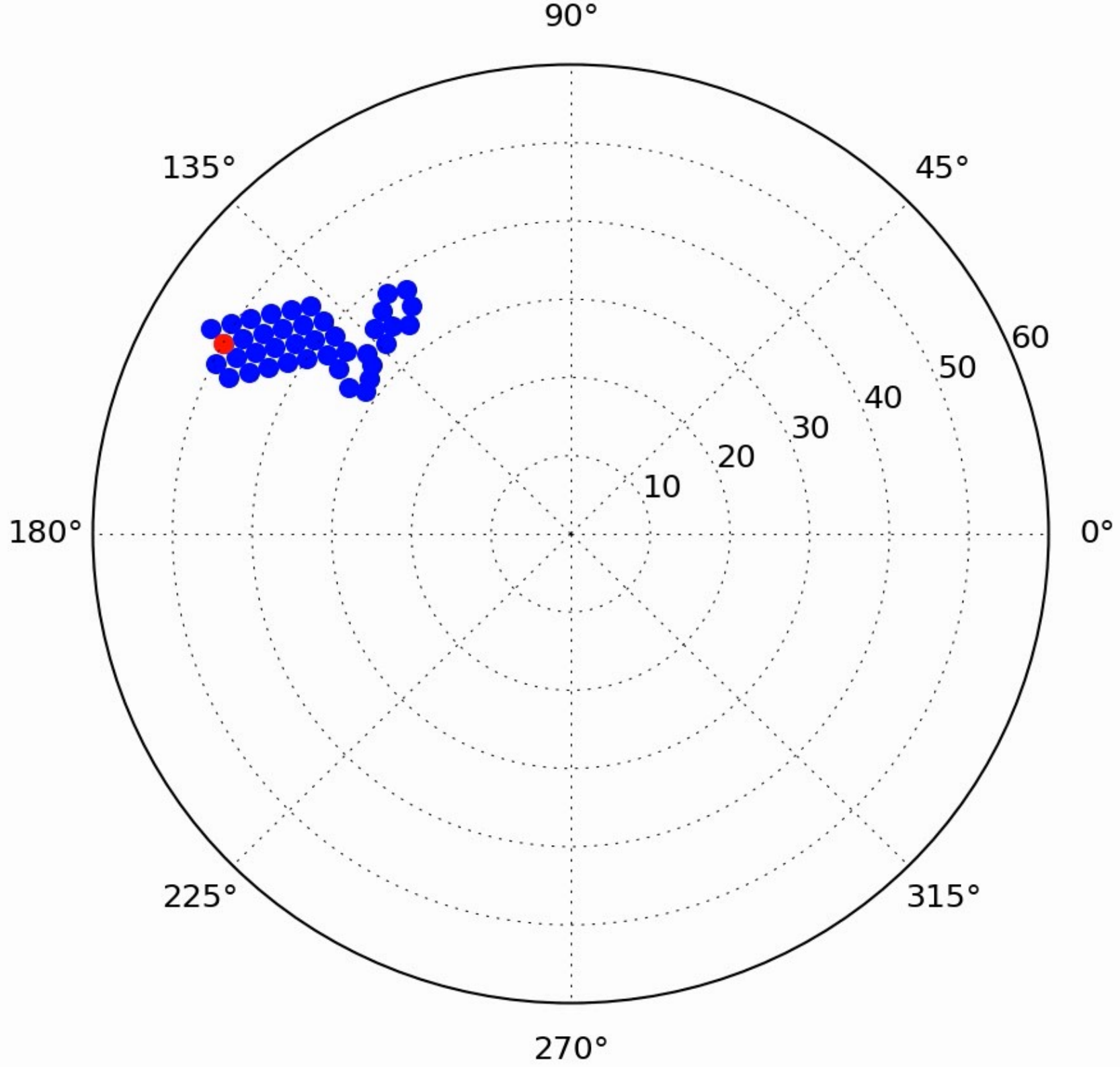


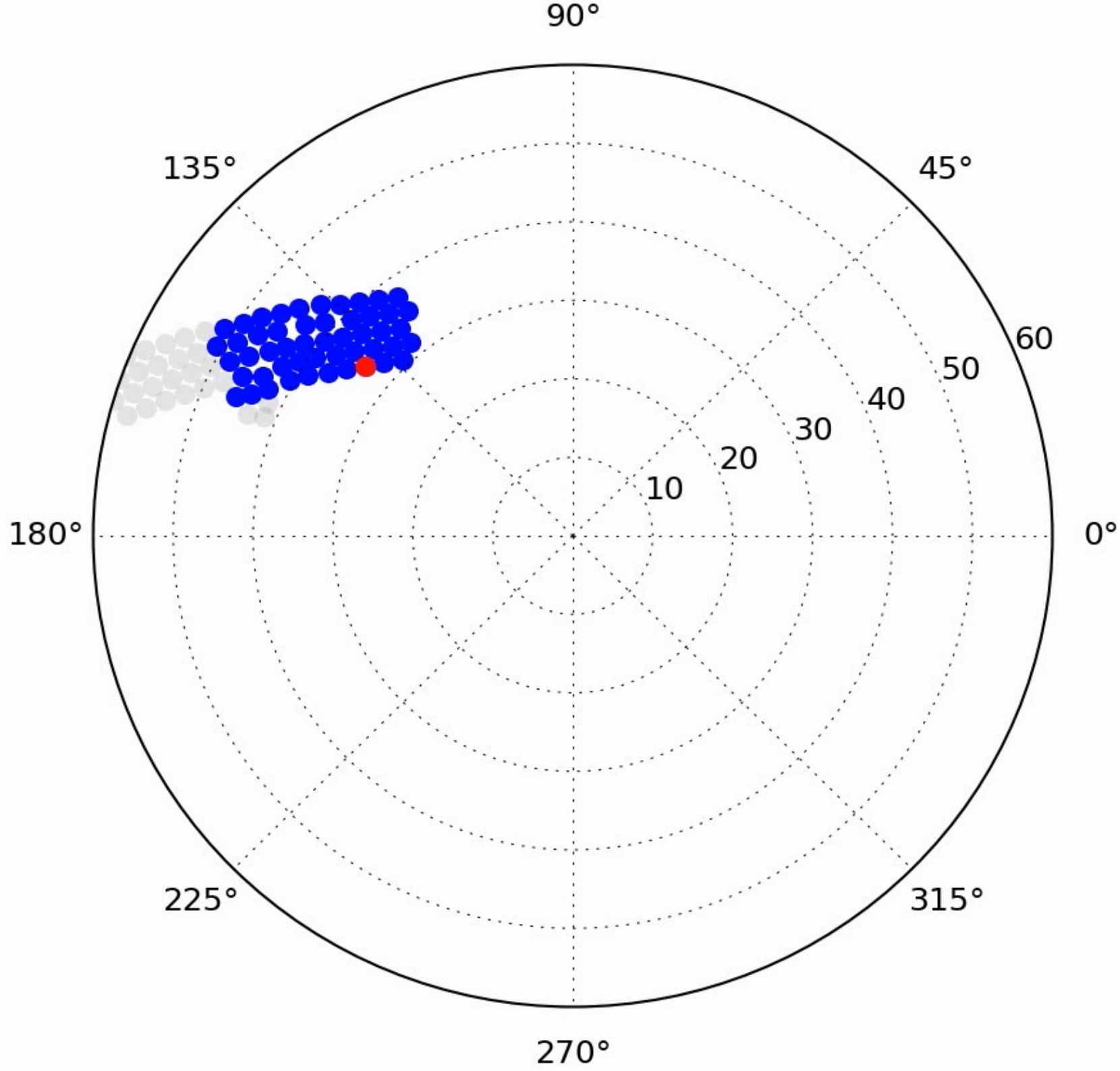
Transient sky



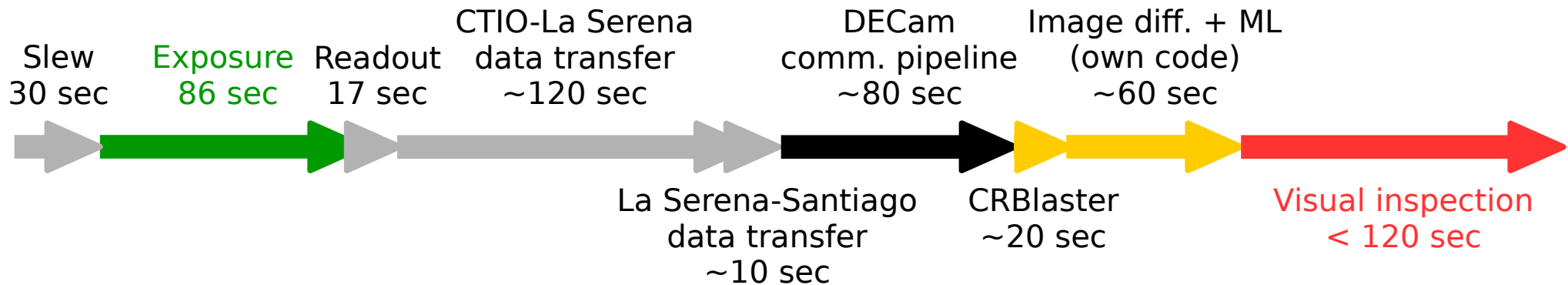
HiTS: searching for shock breakouts in the optical







Pipeline flow outline



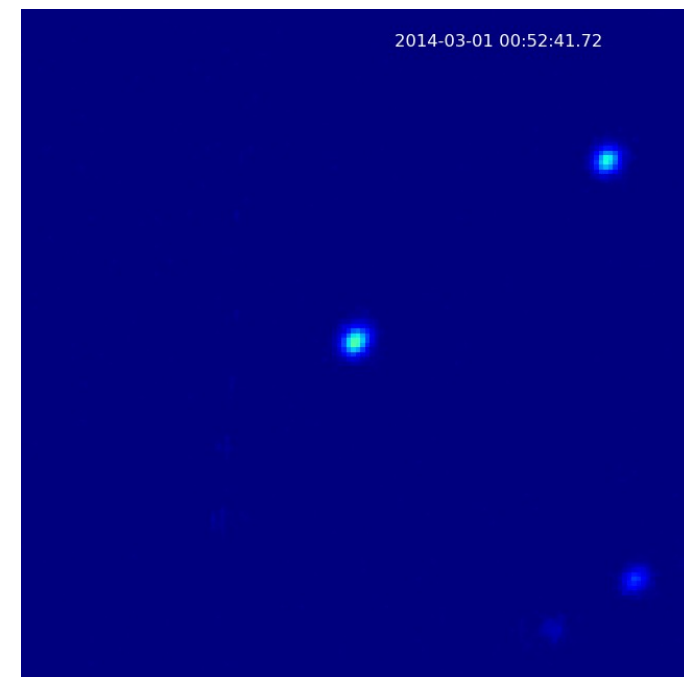
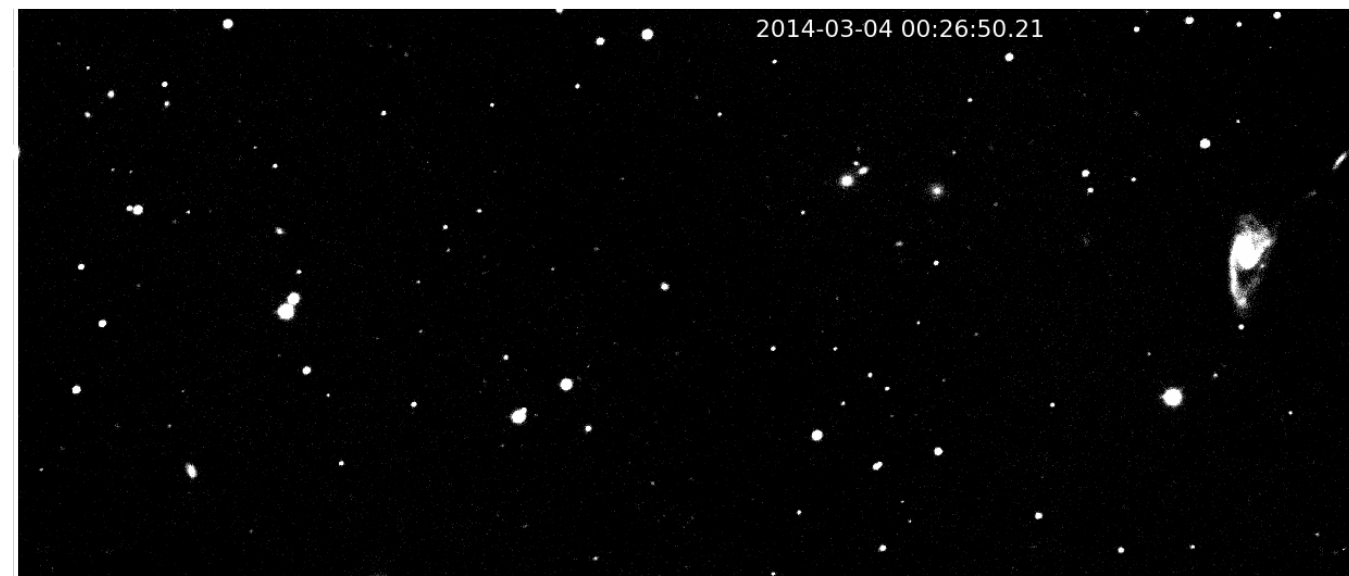
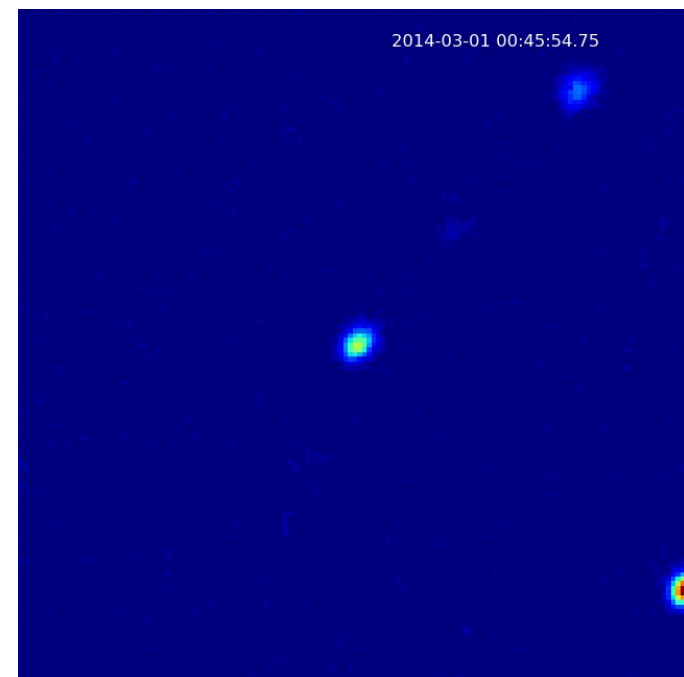
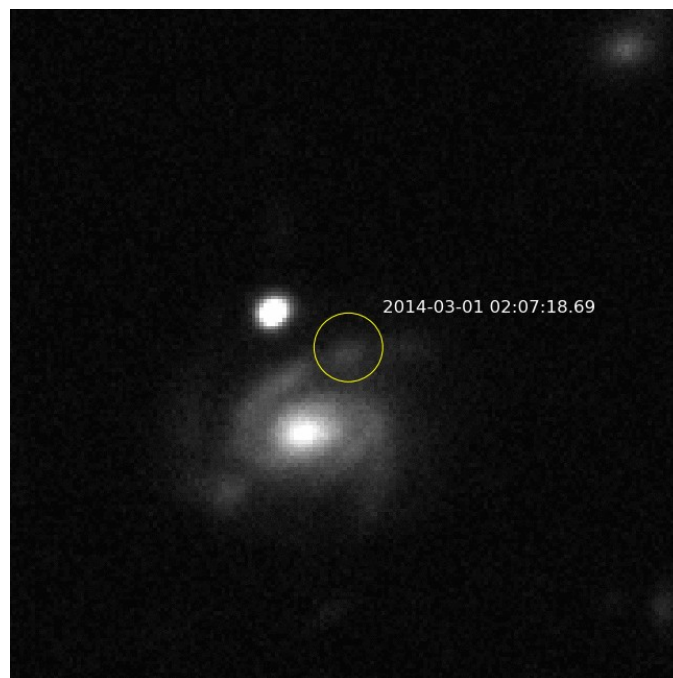
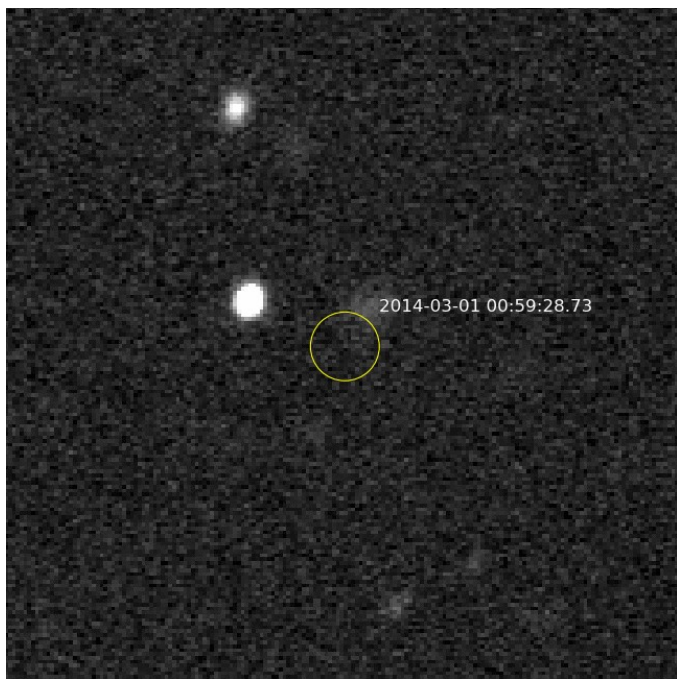
~5-6 min lag



Data stream: 44 Mbps, 4.5 Mpix/s

$\sim 10^{12}$ pixels, $\sim 10^8$ candidates, $\sim 10^6$ filtered candidates

$\sim 10^4$ visual inspections, 125 SNe



ATEL 5949, 5956

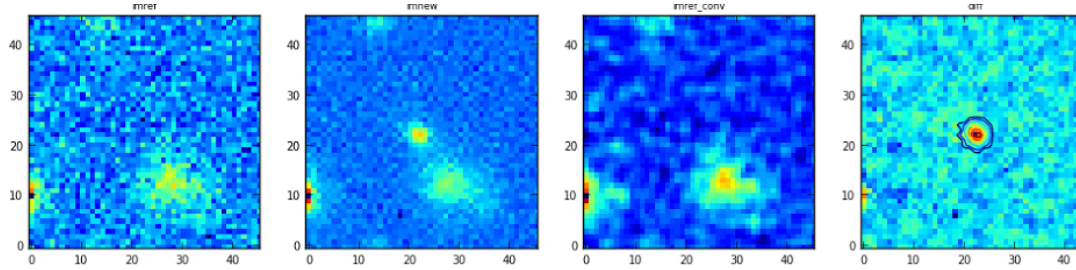
http://www.das.uchile.cl/~fforster/ATEL/summary_das.html

Other science

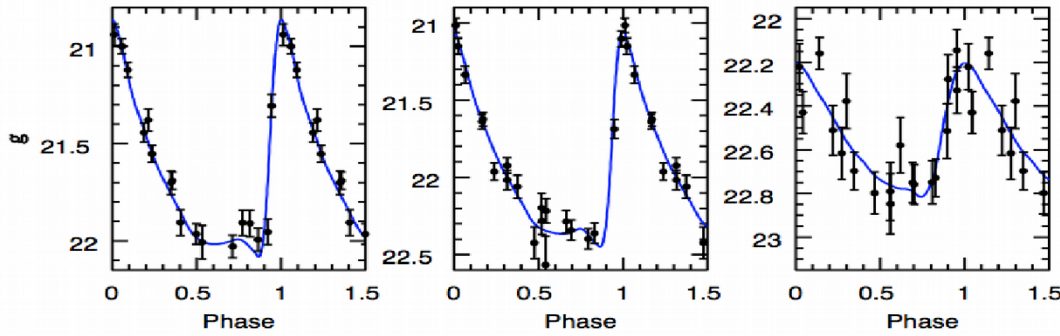
DECam

SOI

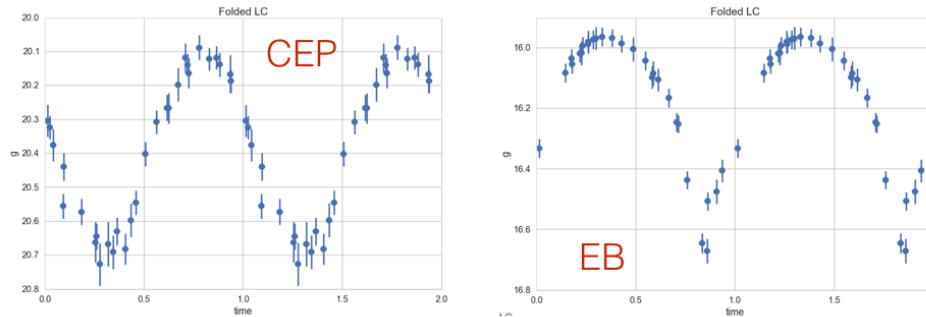
DECam conv. Difference



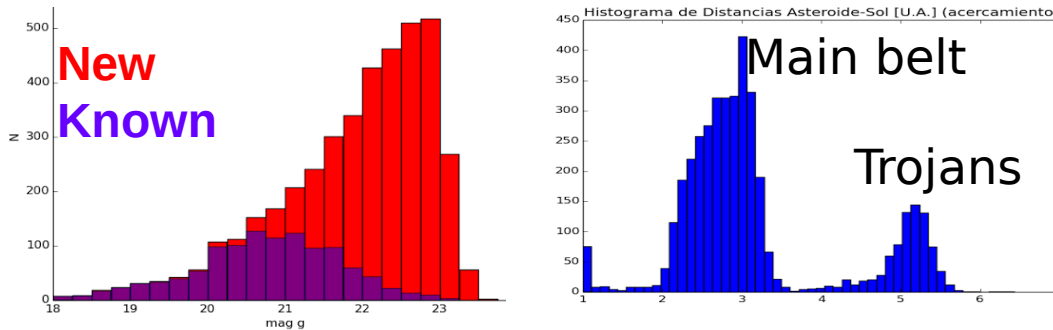
Alessandro Razza:
Combining Decam with
other telescopes



Gustavo Medina:
Most distant RR Lyrae in the
galaxy



Jorge Martínez:
Classifying every transient
object, astrophysics



José Peña
Detecting and
characterizing asteroids



Quimal archiving system



Juan Carlos Maureira

Quimal Archiving System::Dashboard

Logged as: jcm
logout

My Account
My Datasets
Data Explorer
My Shares

Data Explorer

Name	Permissions	Modified	Size	Kind
.size	read and write	Sep 14, 2016 06:12 PM	13 b	Plain text
#README#	read and write	Jul 12, 2013 03:37 PM	66 KB	Plain text
DECAM_00185668.fits.gz	read and write	Mar 12, 2013 10:32 PM	278.40 MB	Application
DECAM_00185672.fits.gz	read and write	Mar 12, 2013 10:38 PM	271.48 MB	Application
DECAM_00185675.fits.gz	read and write	Mar 12, 2013 10:41 PM	271.38 MB	Application
DECAM_00185676.fits.gz	read and write	Mar 12, 2013 10:42 PM	271.41 MB	Application
DECAM_00185677.fits.gz	read and write	Mar 12, 2013 10:44 PM	271.38 MB	Application
DECAM_00185680.fits.gz	read and write	Mar 12, 2013 10:51 PM	271.40 MB	Application
DECAM_00185681.fits.gz	read and write	Mar 12, 2013 10:57 PM	646.30 MB	Application
DECAM_00185683.fits.gz	read and write	Mar 12, 2013 11:02 PM	646.18 MB	Application
DECAM_00185684.fits.gz	read and write	Mar 12, 2013 11:04 PM	646.08 MB	Application
DECAM_00185685.fits.gz	read and write	Mar 12, 2013 11:06 PM	646.13 MB	Application
DECAM_00185688.fits.gz	read and write	Mar 12, 2013 11:13 PM	645.88 MB	Application
DECAM_00185696.fits.gz	read and write	Mar 12, 2013 11:25 PM	434.77 MB	Application
DECAM_00185697.fits.gz	read and write	Mar 12, 2013 11:27 PM	425.51 MB	Application

decam/DECAM/DECAM-13A
Items: 388, size: 124.30 GB

User Identified by: DC=University of Chile, O=Center for Mathematical Modeling, OU=Quimal Project, CN=Juan Carlos Maureira Bravo
Portal ONLINE

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Center for Mathematical Modeling - University of Chile

Dataset



Files



Volumes



Tapes



LTO6 - 2.5 TB raw, 6.25 TB 2:5:1 compression

LTFS - Linear Tape File system

PKI - Public Key Infrastructure

LTFS
Linear Tape File System



ULTRIUM 6
LTO
- 2.5TB native 6.25TB compressed
- 160MB/s native 400MB/s compressed



National Laboratory for High Performance Computing (NLHPC)



Fastest computer in Chile, 2nd
in Latin America (LARTop50)
2640 cores (44 TFlops)
Infiniband (FDR 56 Gbps)
Lustre storage (DDN)

Services for:
Academia
Public sector
Industry

Team of HPC experts

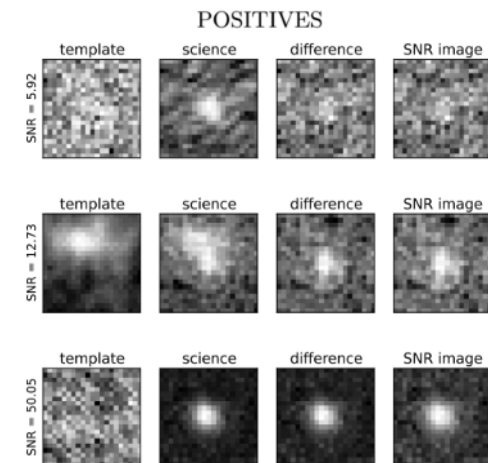
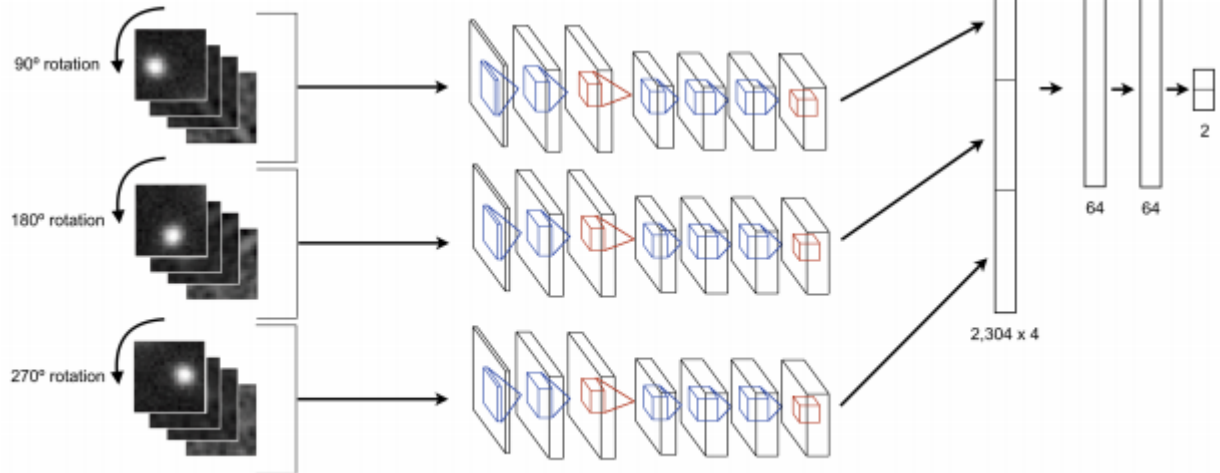
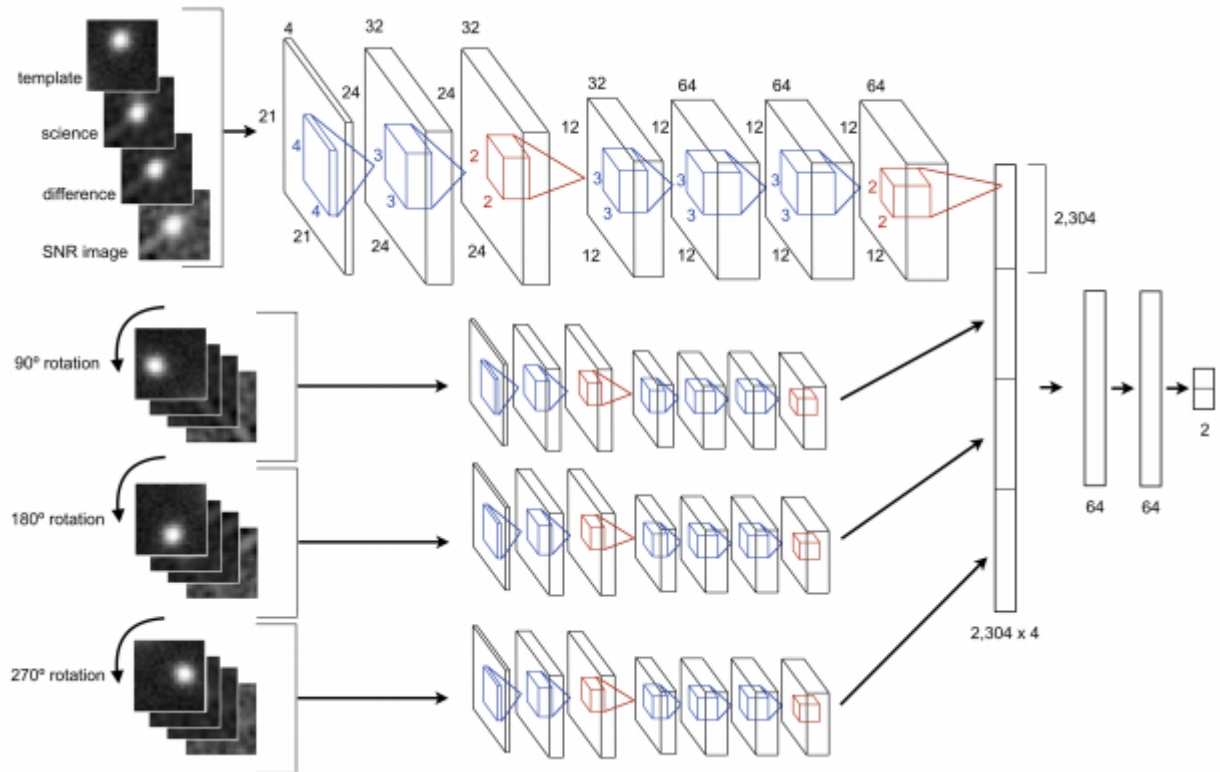
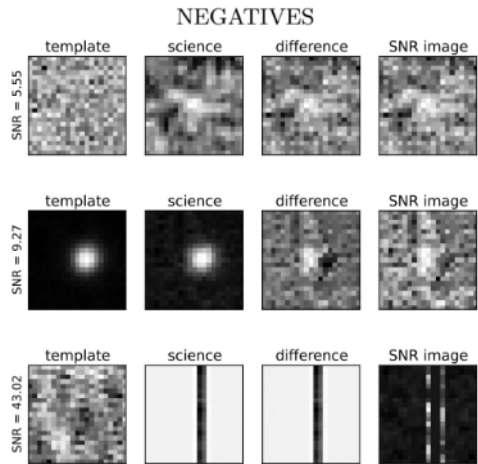
Deep learning for real bogus separation



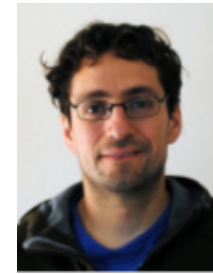
Ignacio Reyes
DIE



Guillermo Cabrera
CMM, U. Chile



Un-biased Supervised Learning



Guillermo
Cabrera
CMM, U. Chile



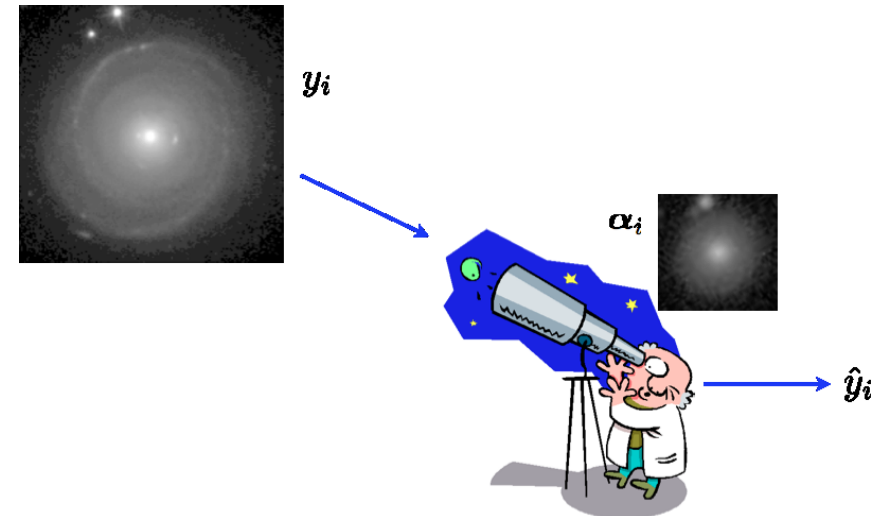
Chris Miller
Astronomy, U.
Michigan



Jeff Schneider
Robotics Institute, CMU



- Biased labels due to observable parameters
- Quantify the labeling bias.
- De-bias the labels.



Interferometric Image Reconstruction

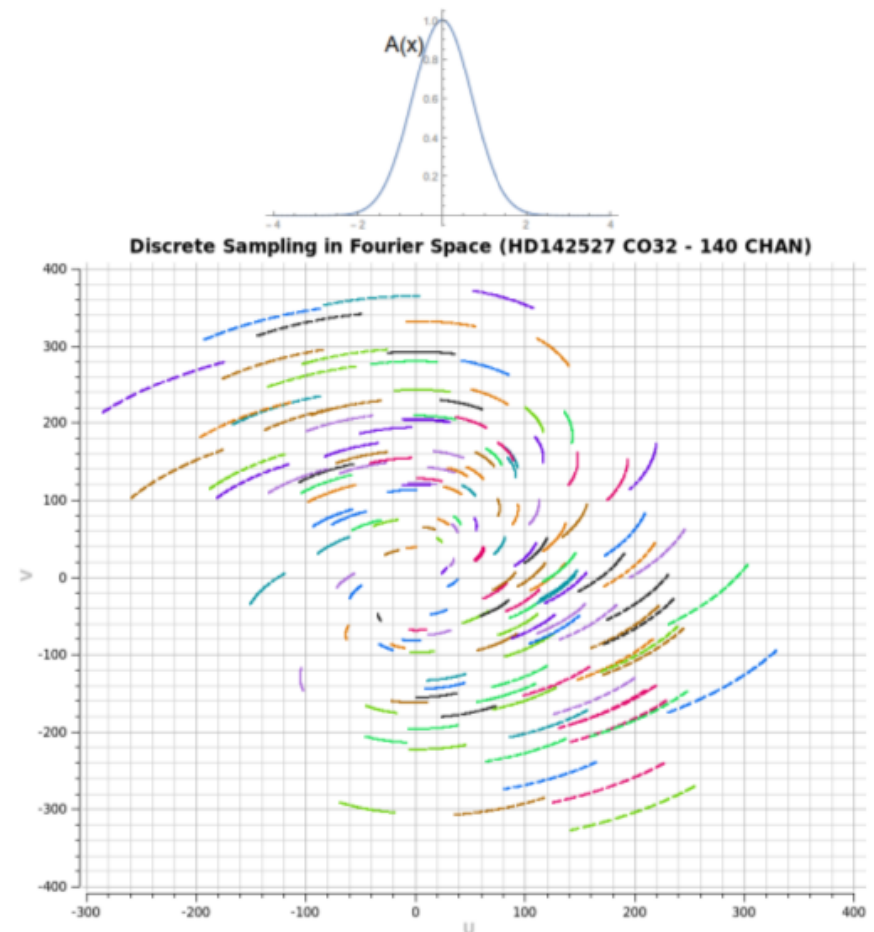


Pablo Román
Postdoc
Ph.D. Systems Eng.



Simón Casassus
Professor
Ph.D. Astrophysics

- 1 $V(u, v) = \mathcal{F}[\mathcal{A} \cdot I](u, v)$
- 2 V function is a random variable: is a Random Field.
- 3 We want: estimation for $E(I)$.
- 4 We want: estimation for $E(I^2)$.
- 5 Our data: (Simplified version)
 - Sampling points $\{z_k = (u_k, v_k)\}_{k=1}^N$,
 $N \sim 10^4 - 10^6$
 - Sampled function values at those points $\{V_k^o\}_{k=1}^N$
 - Variance estimation of measurement $\{\sigma_k\}_{k=1}^N$



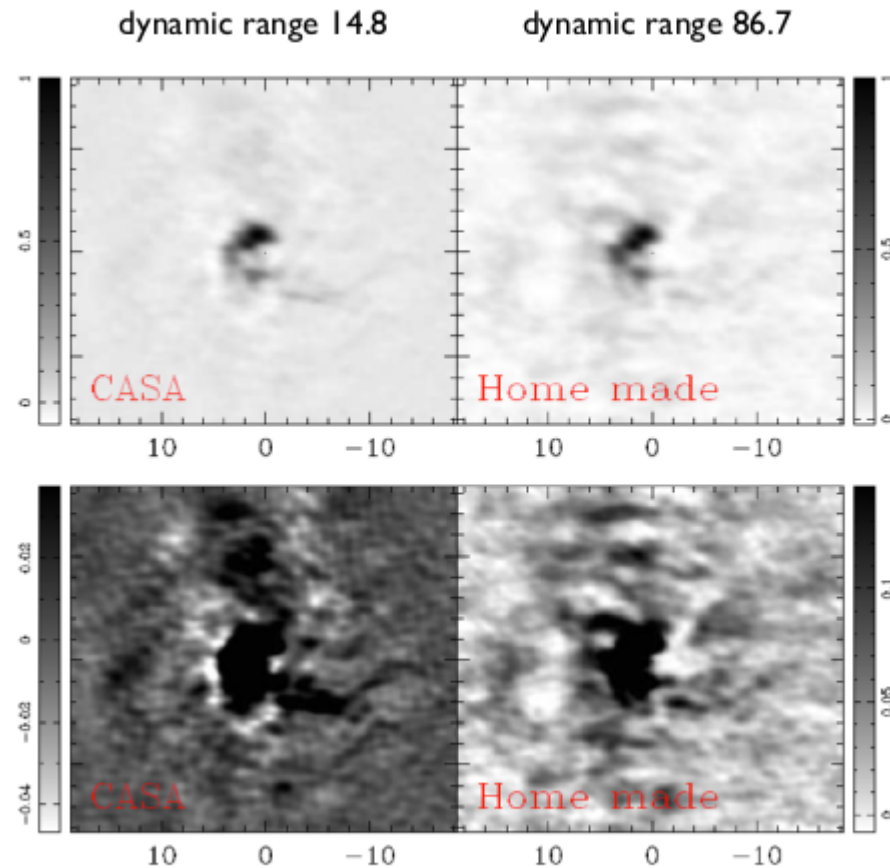
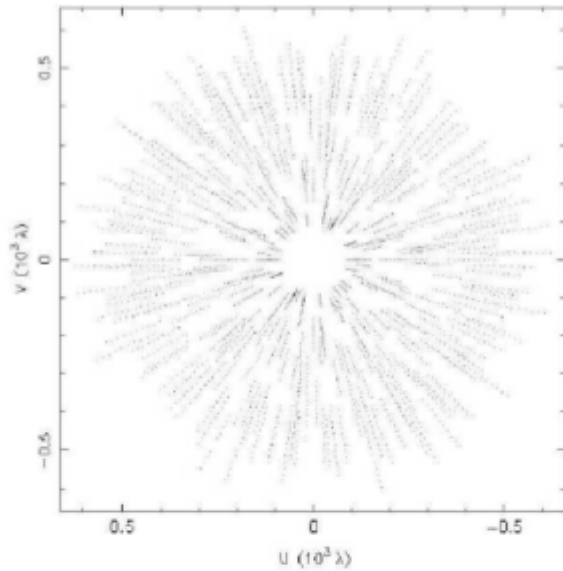
Interferometric Image Reconstruction



Pablo Román
Postdoc
Ph.D. Systems Eng.

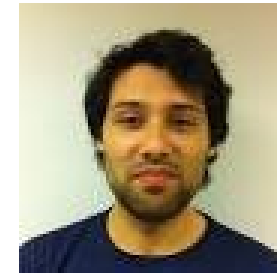


Simón Casassus
Professor
Ph.D. Astrophysics



dynamic range: the ratio of the peak signal over the norm of the most negative artifact.

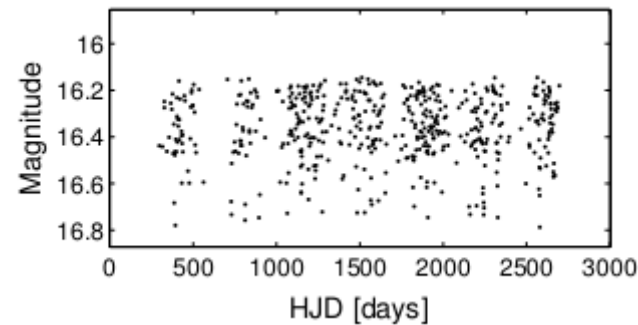
Correntropy Non-negative Matrix Factorization (NMF) Spectrum



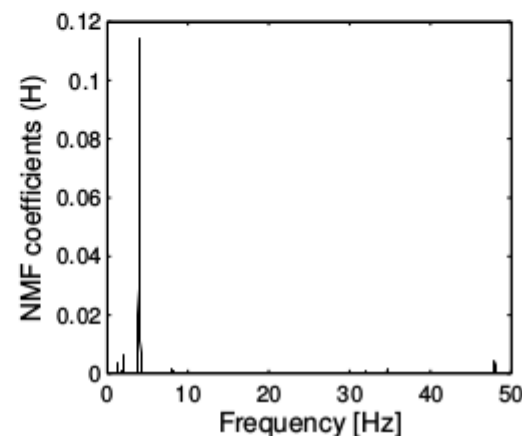
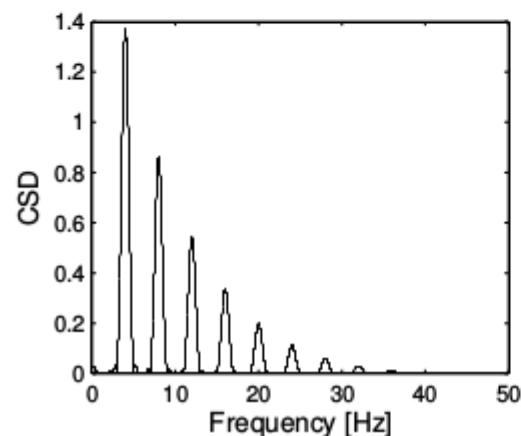
Pablo Huijse
Postdoc
Ph.D. Elec. Eng.



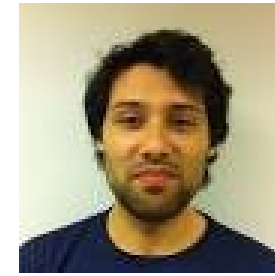
Pablo Estévez
Professor
Ph.D. Elec. Eng.



- Decompose correntropy into a frequency dictionary
- The NMF coefficients are the spectral ordinates of the periodogram
- Overcomplete dictionary + optimization procedure: sparse and high-res periodogram
- Adapt the kernel size of the atoms to further increase resolution



Correntropy NMF Spectrum



Pablo Huijse
Postdoc
Ph.D. Elec. Eng.



Pablo Estévez
Professor
Ph.D. Elec. Eng.

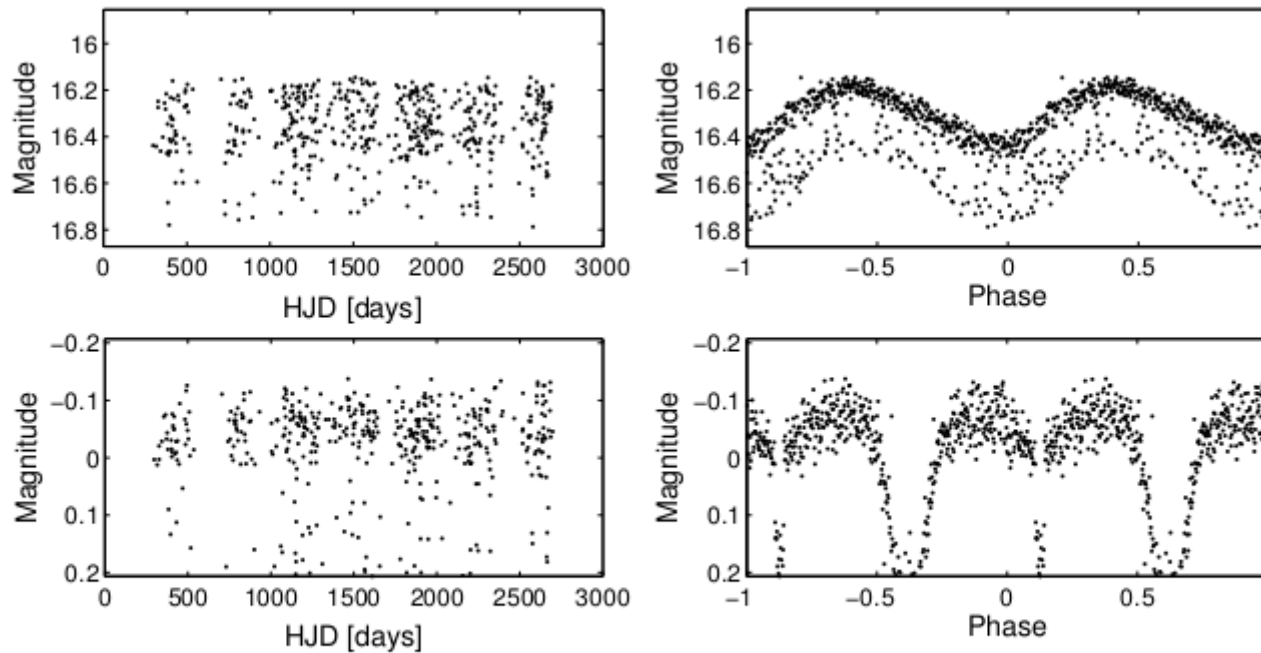


Figure: Pulsating with a 3.6 days periods plus EB with 68 days period.

The future

- Big opportunities ahead for Chilean astronomy, but bigger challenges for traditional astronomy → growth of astroinformatics
- Data–driven, interdisciplinary **teams** needed.
- Access to the National Laboratory for High Performance Computing a unique opportunity
- Astroinformatics must tackle real astrophysical questions and work closely with astronomers to be relevant
- Access to large volumes of data is a test bed for new data analysis tools with potential applications in other fields (e.g. ALMA image synthesis and medical imaging, LSST for real–time data stream analysis, image processing, machine learning)
- CMM among first truly interdisciplinary centers in the world, pioneered data–driven science in Chile. **Focus on technology transfer.**

New York Times

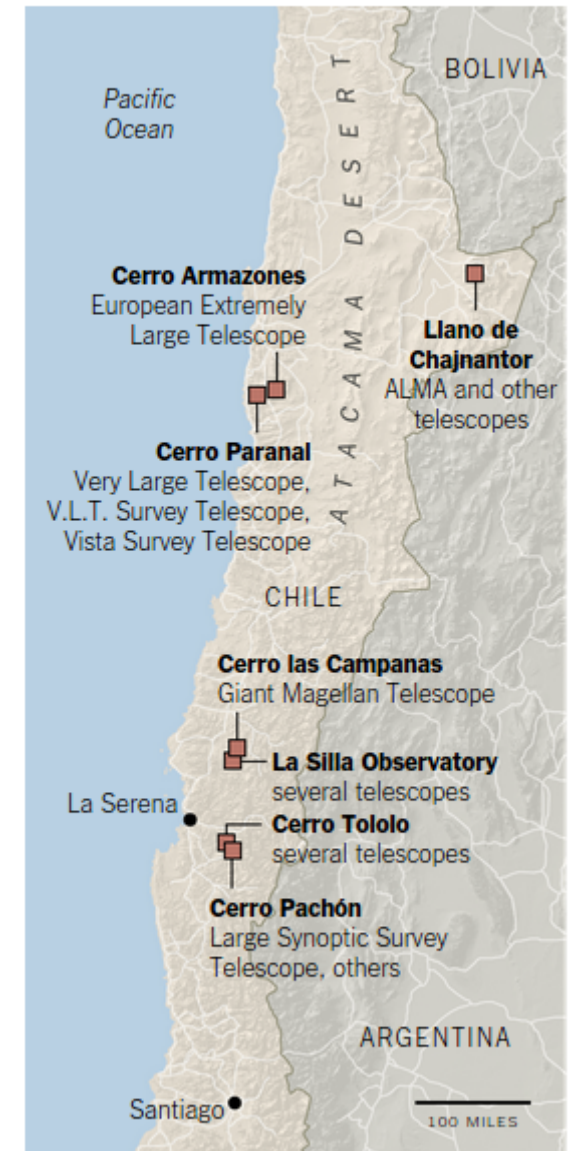
“More Eyes on the Skies”, July 2014

“If it all plays out as expected and budgeted, astronomers of the 2020s will be swimming in petabytes of data streaming from space and the ground”

“The inauguration of these new telescopes, early in the next decade, will further enshrine the Atacama Desert in Chile, which is bone-dry, high, dark and blessed with remarkably steady air, as the center of world astronomy”

Looking Up

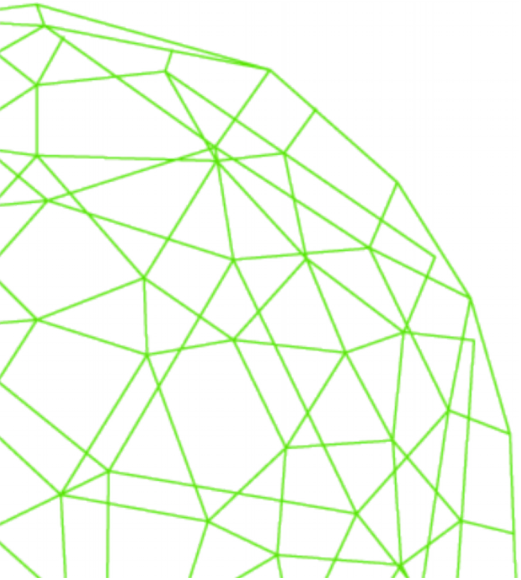
Northern Chile is dry and high, with little light pollution. The area already bristles with telescopes, and more are under construction.



Source: European Southern Observatory

By The New York Times

THANKS!





The astrophinformatics Laboratory at CMM



Jaime San Martín
Professor
Ph.D. Mathematics



Eduardo Vera
Scientist
Ph.D. Physics



Guillermo Cabrera
Engineer
Ph.D. Comp. Sc.



Pablo Román
Postdoc
Ph.D. Systems Eng.



Francisco Förster
Scientist
Ph.D. Astronomy



J. Carlos Maureira
Scientist
Ph.D. Comp. Sc.



Santiago González-Gaitán
Postdoc
Ph. Astronomy

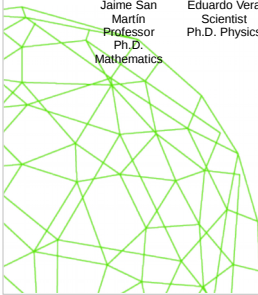


Pablo Huijse
Postdoc
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Pablo Estévez
Professor
Ph.D. Elec. Eng.

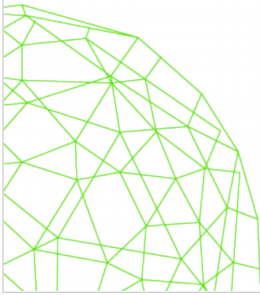
and many more collaborators...



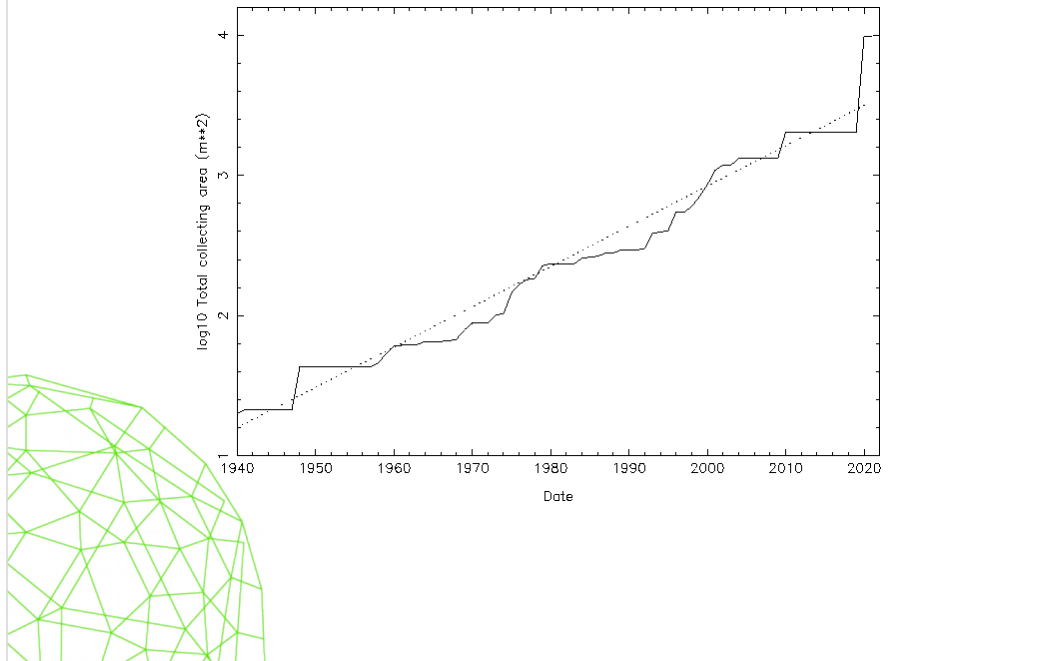
Astroinformatics Laboratory

- Focus in astronomical problems that require the analysis of **large volumes of data**.
- Efficient algorithms, statistics, data handling methods, visualization, inverse problems, image processing, machine learning, high performance computing.
- **Interdisciplinary** focus gives **flexibility** to move into new developing areas of astronomy.
- Focus on **data-driven** and well defined problems.
- e.g. real-time, high-cadence survey for transient objects (**DECam**); automatic classification of galaxies (**LSST**), characterization of light curves (**DECam**, **LSST**) and image synthesis in interferometric observations (**ALMA**).

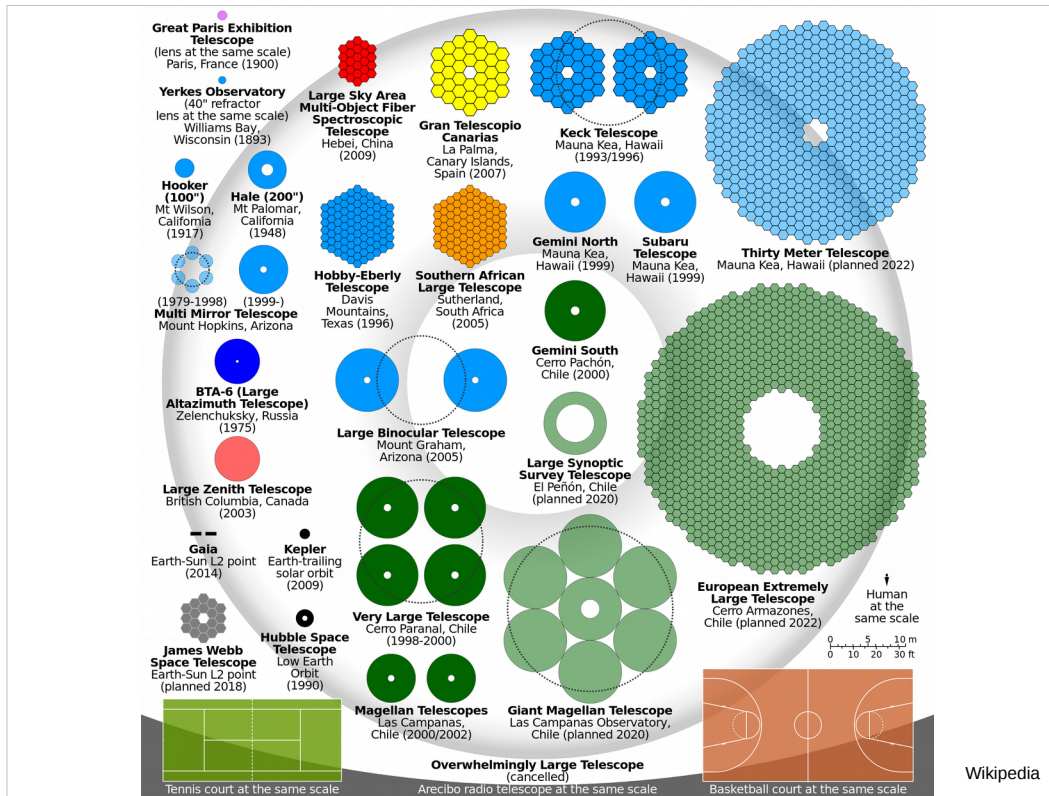
**Astronomy in the next decade:
Big Data & Big Telescopes**



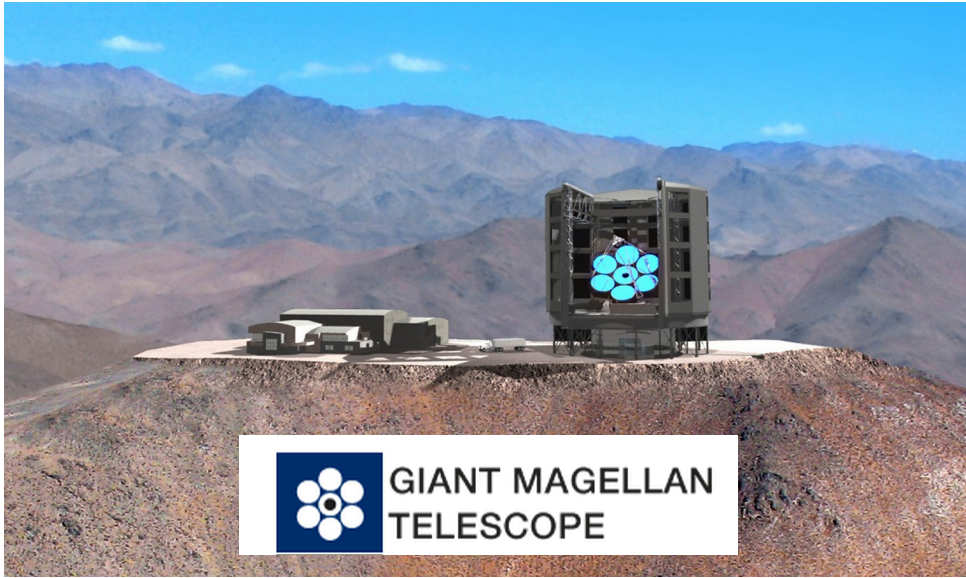
Big Telescopes



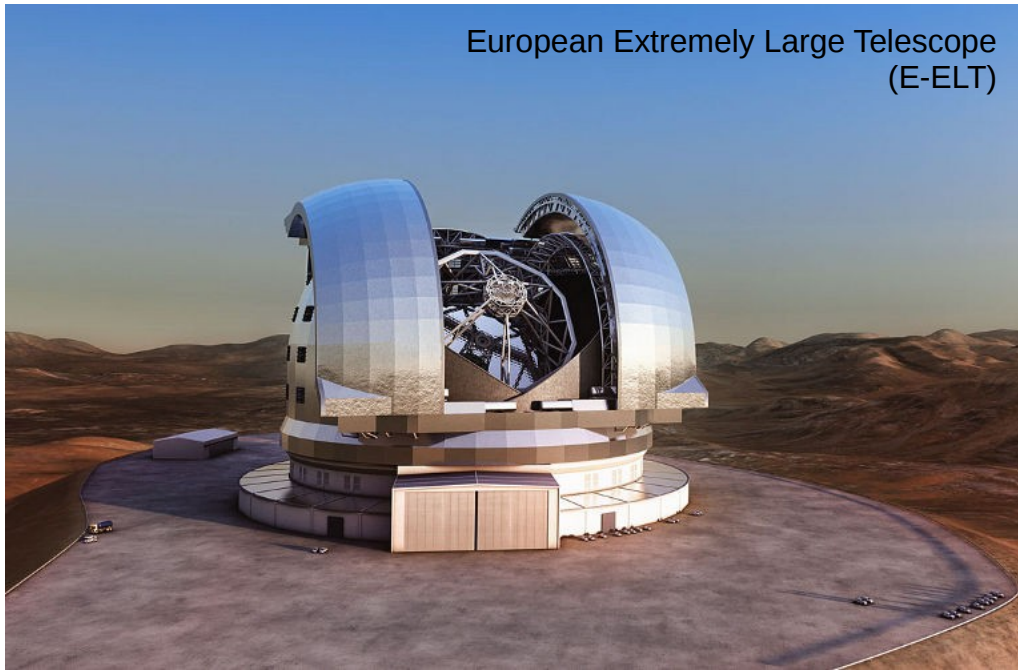
1era revolución en astronomía: evolución exponencial del área colectora de los telescopios



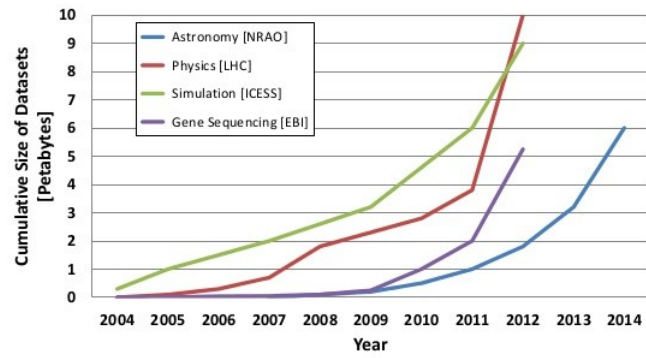
Tamaños de telescopios. En verde son telescopios instalados o a instalarse (verde claro) en Chile.



European Extremely Large Telescope
(E-ELT)



Big Data in astronomy



Scientific Data Grows Exponentially!

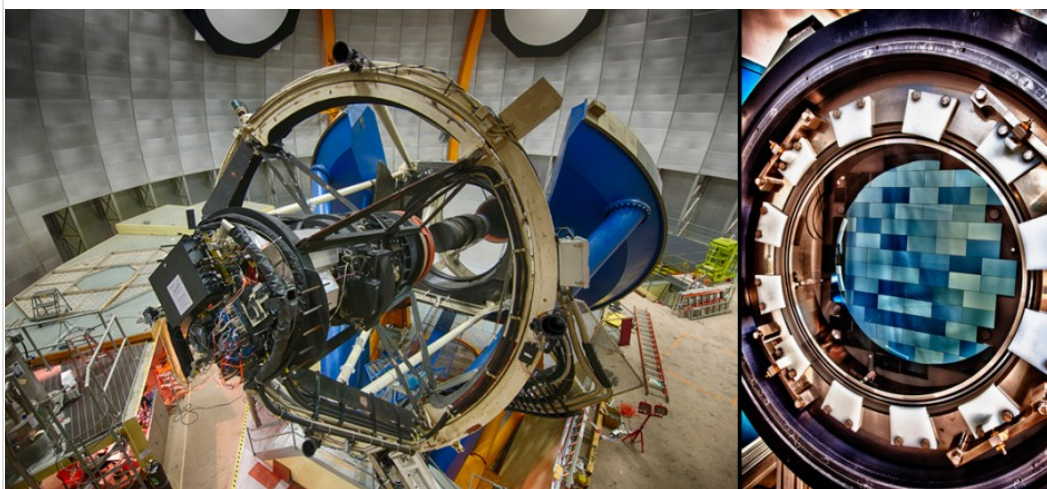
3

Thomas Heinis, Human Brain Project

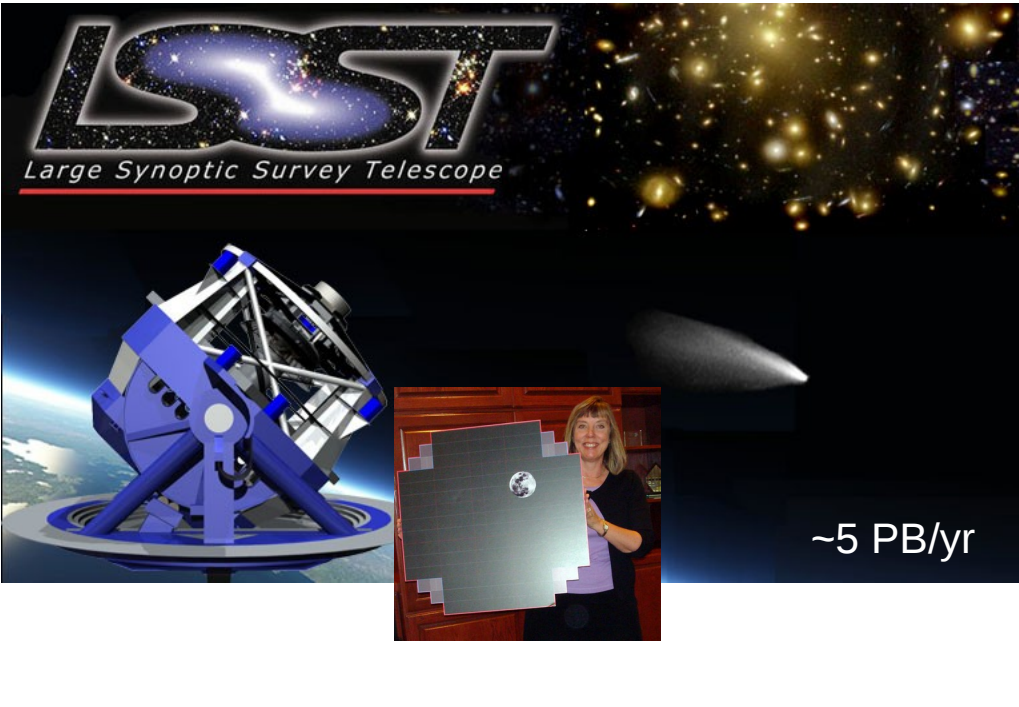


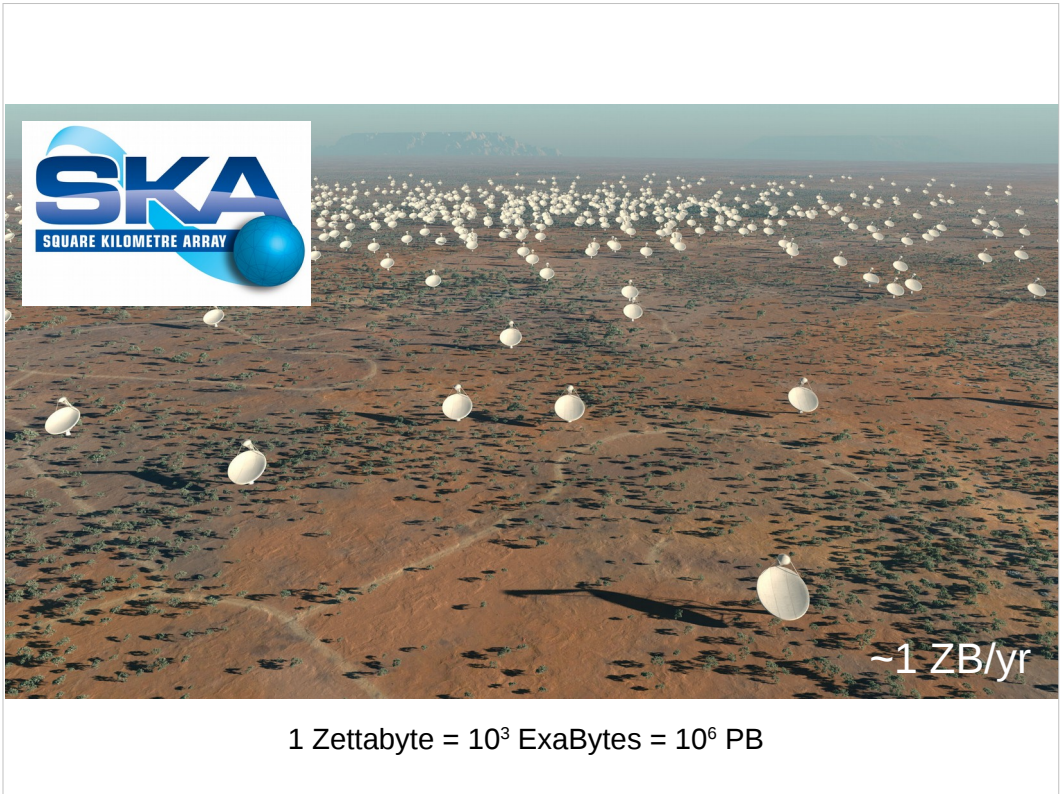
~0.2 PB/yr

Dark Energy Camera (DECam)



~0.4 PB/yr



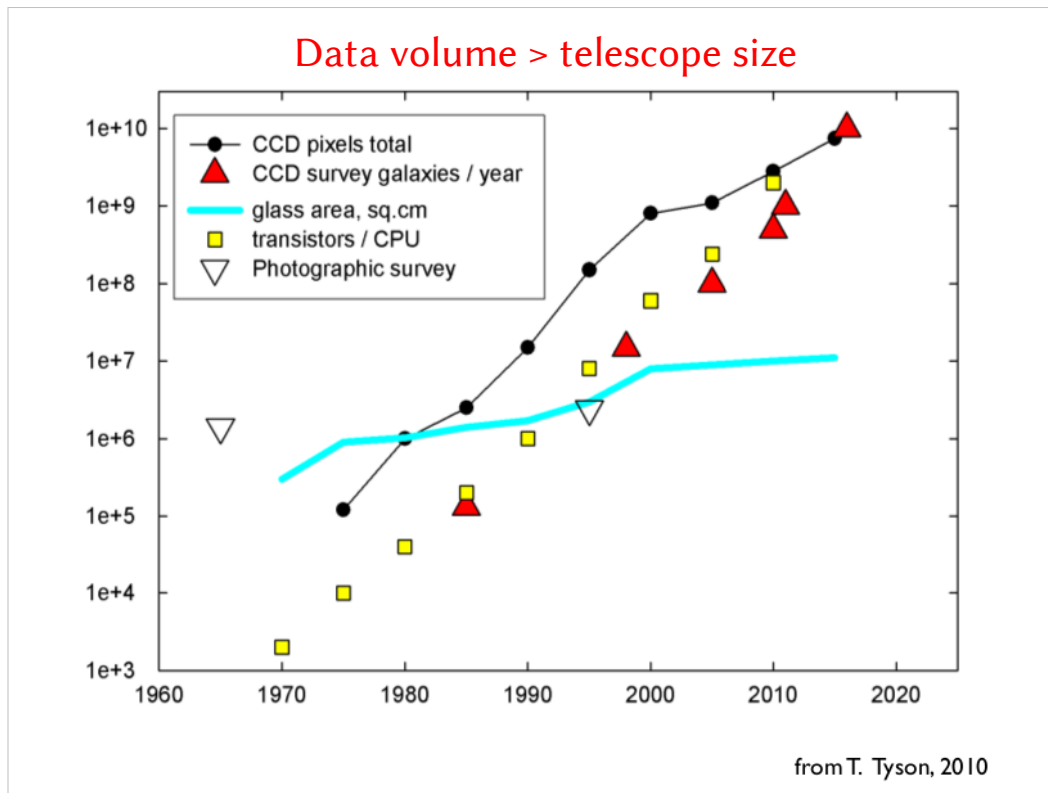


~1 ZB/yr

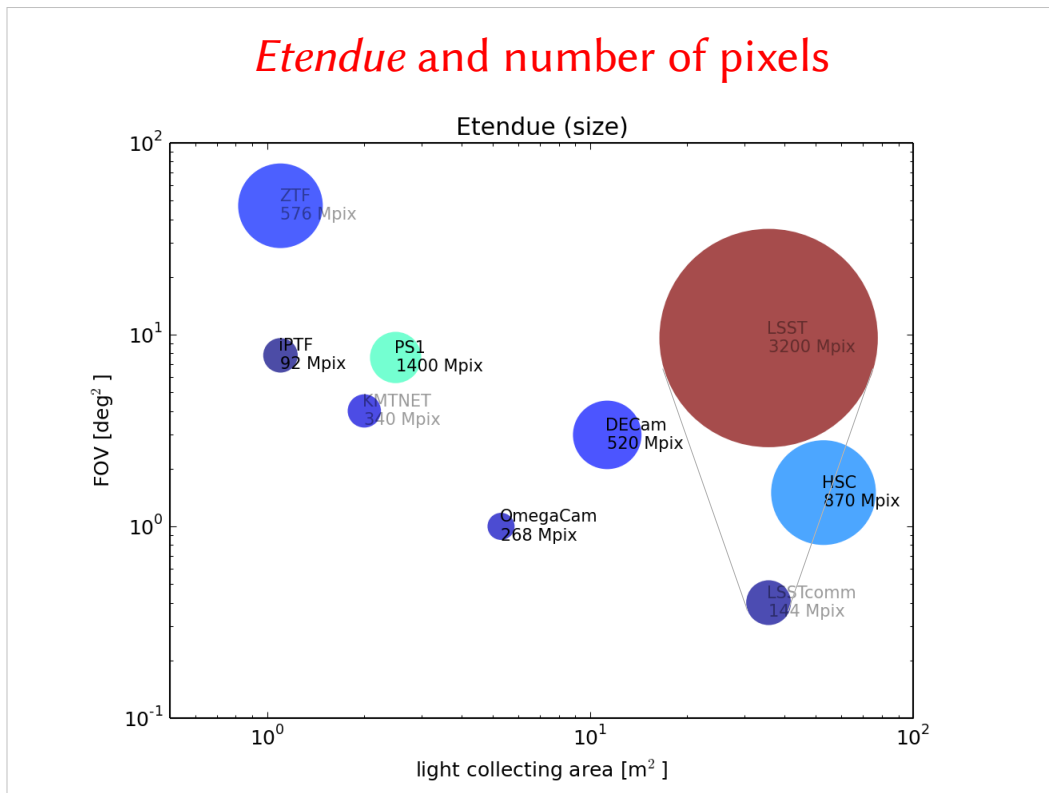
1 Zettabyte = 10^3 ExaBytes = 10^6 PB



El problema de cómo almacenar los datos de SKA no está resuelto. Hacerlo en la nube es muy caro, la solución probablemente sea distribuida



2da revolución en astronomía: ley de Moore se ve reflejada en el número de galaxias observadas



Área en el cielo vs área del espejo primario de diferentes telescopios. Producto de ambos se conoce como etendue y da cuenta de la capacidad de observar un mayor volumen del universo en una sola toma.

The large Synoptic Survey Telescope (LSST)

Time domain science

- Nova, supernova, GRBs
- Source characterization
- Instantaneous discovery

Census of the Solar System

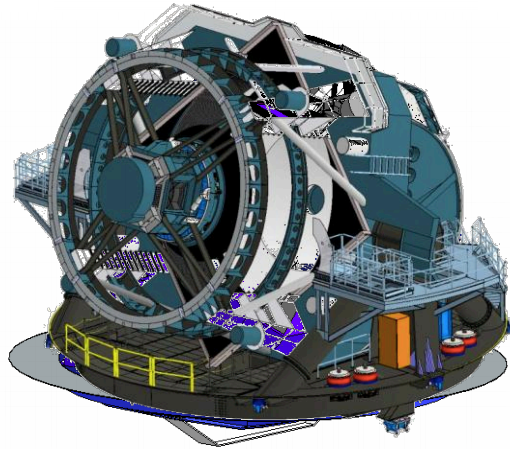
- NEOs, MBAs, Comets
- KBOs, Oort Cloud

Mapping the Milky Way

- Tidal streams
- Galactic structure

Dark energy and dark matter

- Strong Lensing
- Weak Lensing
- Constraining the nature of dark energy



The large Synoptic Survey Telescope (LSST)

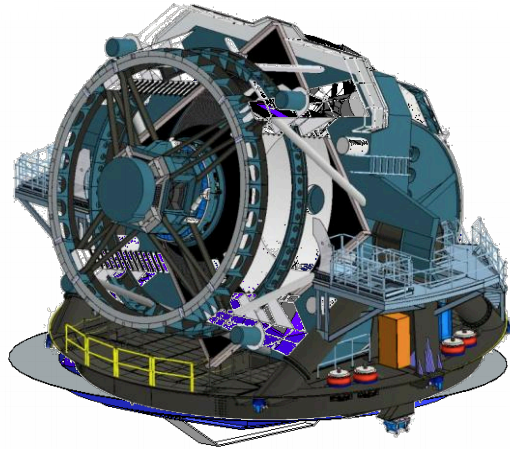
Traditional scientists

What data should I collect to test an hypothesis?

Data-driven scientist

What theories can I test given the data I already have?

Success in research will depend on the ability to mine knowledge from data.



The large Synoptic Survey Telescope (LSST)

LEVEL 1: Nightly alerts

~10 million time-domain events per night

~6 million bodies in the Solar System

LEVEL 2: Yearly data releases

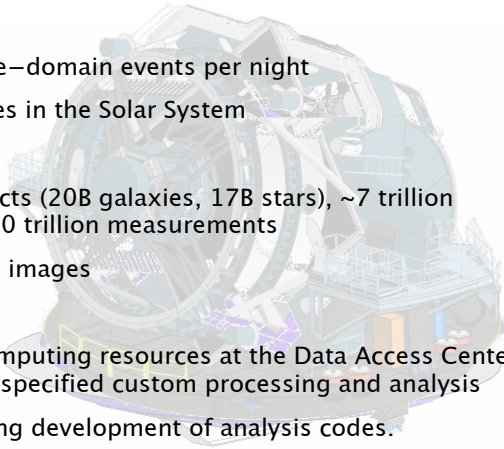
~37 billion objects (20B galaxies, 17B stars), ~7 trillion sources, and ~30 trillion measurements

Deep co-added images

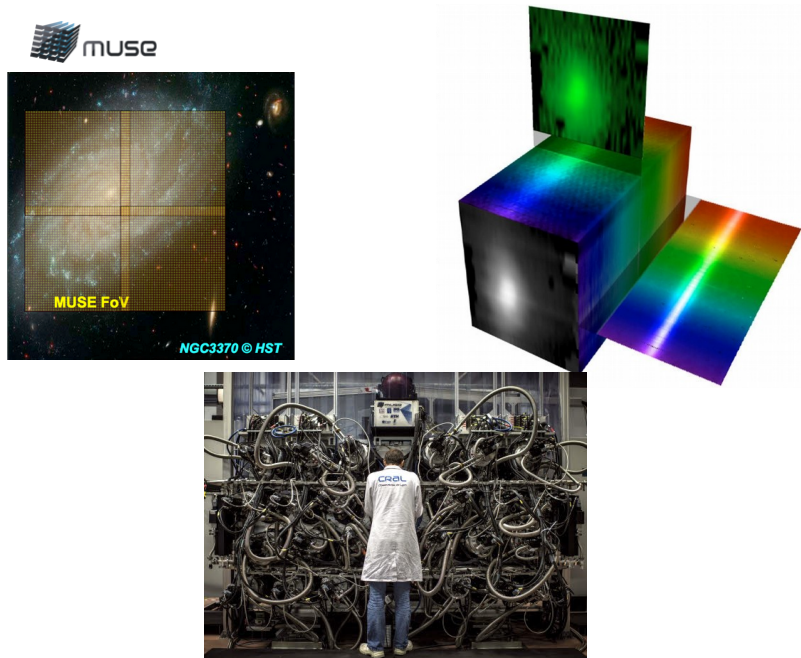
LEVEL 3: Community projects

Services and computing resources at the Data Access Centers to enable user-specified custom processing and analysis

Software enabling development of analysis codes.



Integral Field Units, e.g. Multi-Unit Spectroscopic Explorer (MUSE)



MUSE es un integral field unit telescope, que produce cubos de datos de 10 mil espectros en el óptico con excelente resoluciónn espectral y espacial

Internet submarine cables





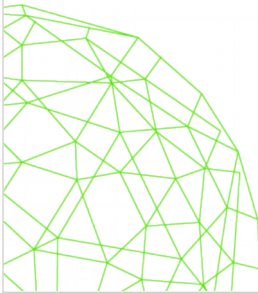
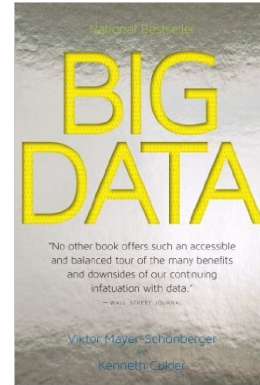
Big data

“Big Data is all about seeing and understanding the relations within and among pieces of information...”

“It is about three major shifts of mindset ... the ability to analyze **vast amounts of data** about a topic ... willingness to embrace **data's real-world messiness** ... growing respect for **correlations** rather than a continuous quest for elusive **causality**.”

“Simple models and a lot of data trump over more elaborate models on less data”

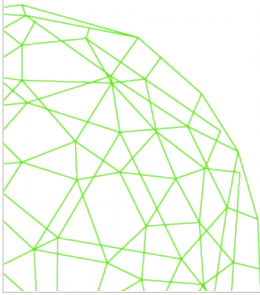
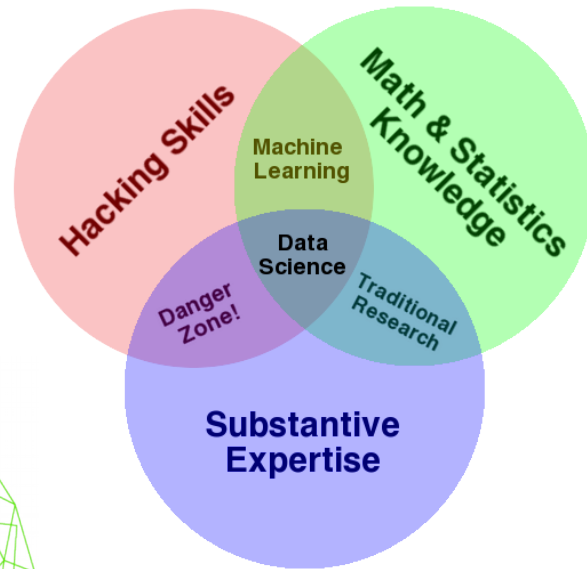
Peter Norvig, Google artificial intelligence.



"Big data is like teenage sex: everyone talks about it, nobody really knows how to do it, everyone thinks everyone else is doing it, so everyone claims they are doing it..."

Dan Ariely,
Duke University

Data science



**Harvard
Business
Review**

2012

DATA

Data Scientist: The Sexiest Job of the 21st Century

by **Thomas H. Davenport** and **D.J. Patil**

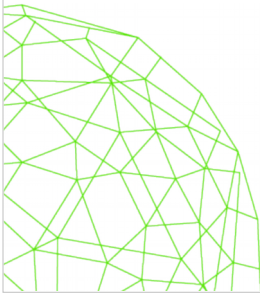
FROM THE OCTOBER 2012 ISSUE

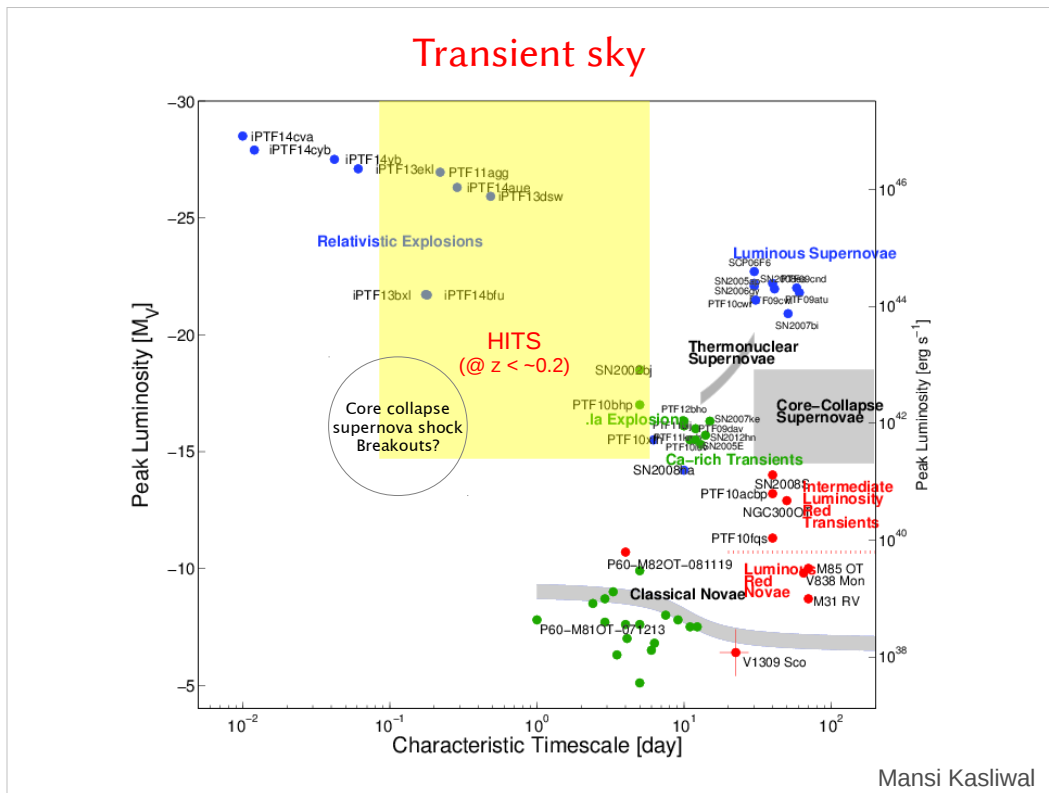
The shortage of data scientists is becoming a serious constraint in some sectors.

Data scientists today are akin to the Wall Street “quants” of the 1980s and 1990s.



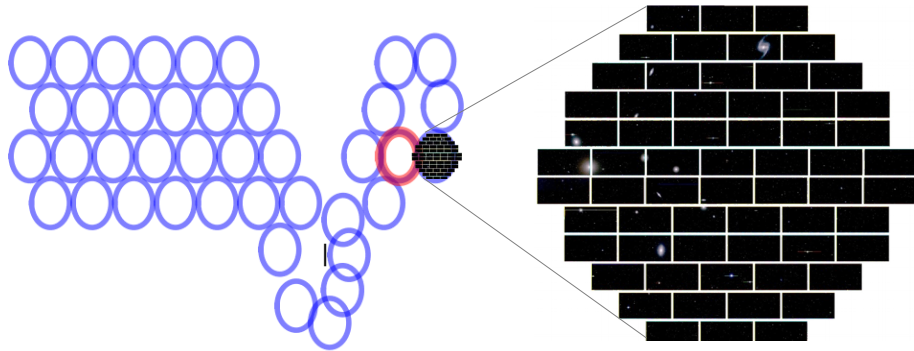
Astroinformatics Laboratory Projects



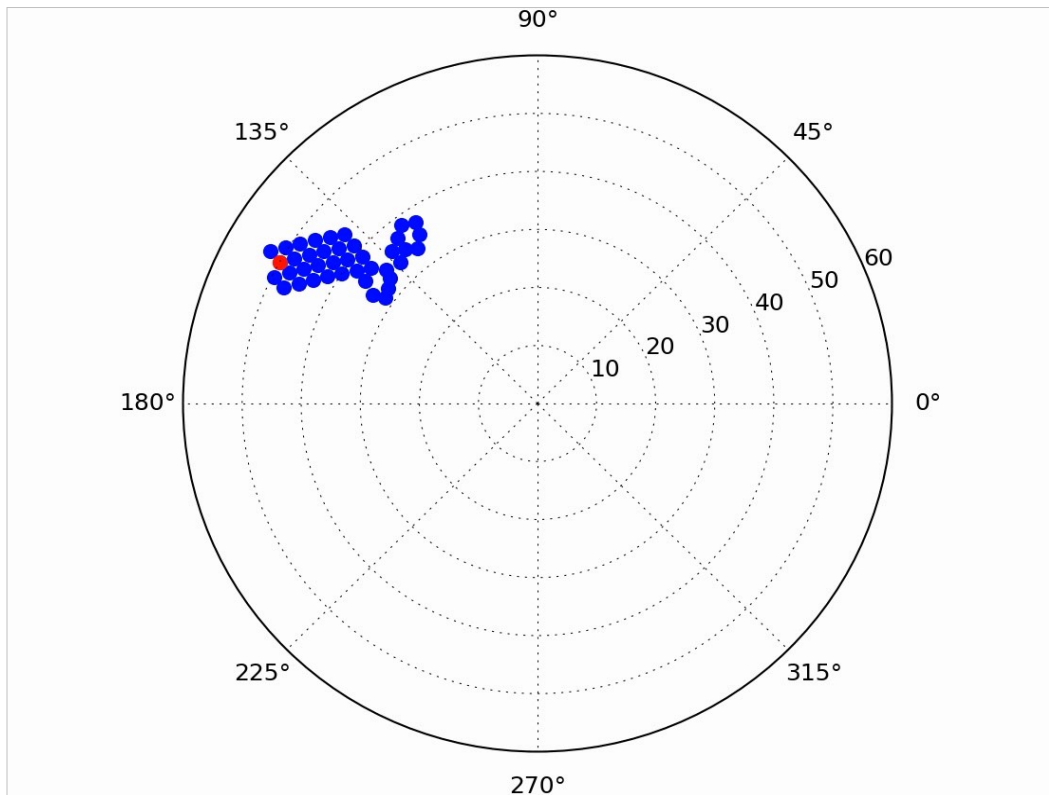


Brillo absoluto vs escala de tiempo característica de diferentes clases de transientes. Región de arriba a la izquierda corresponde a afterglows ópticos de gamma ray bursts (formación de agujeros negros), región de abajo a la derecha corresponde a explosiones estelares de la mayoría de la estrella (supernovas) o parciales (novas). Región de abajo a la izquierda del diagrama está despoblada por la dificultad técnica de las observaciones y es la zona que queremos estudiar con HiTS.

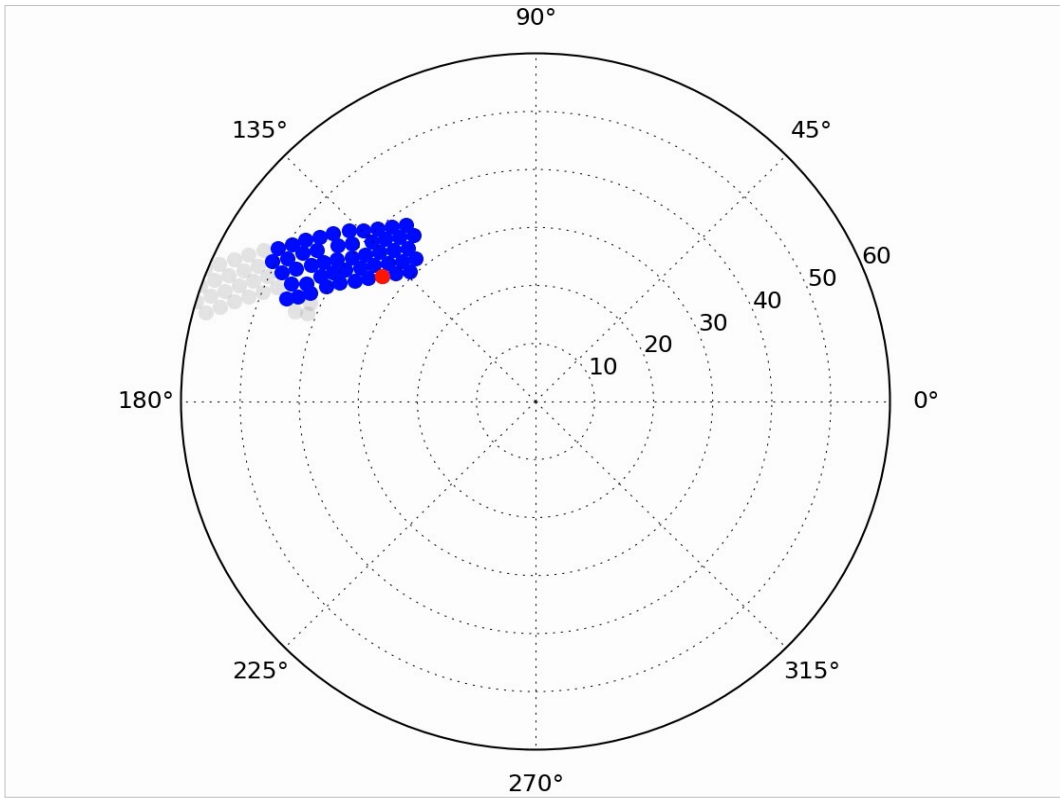
HiTS: searching for shock breakouts in the optical



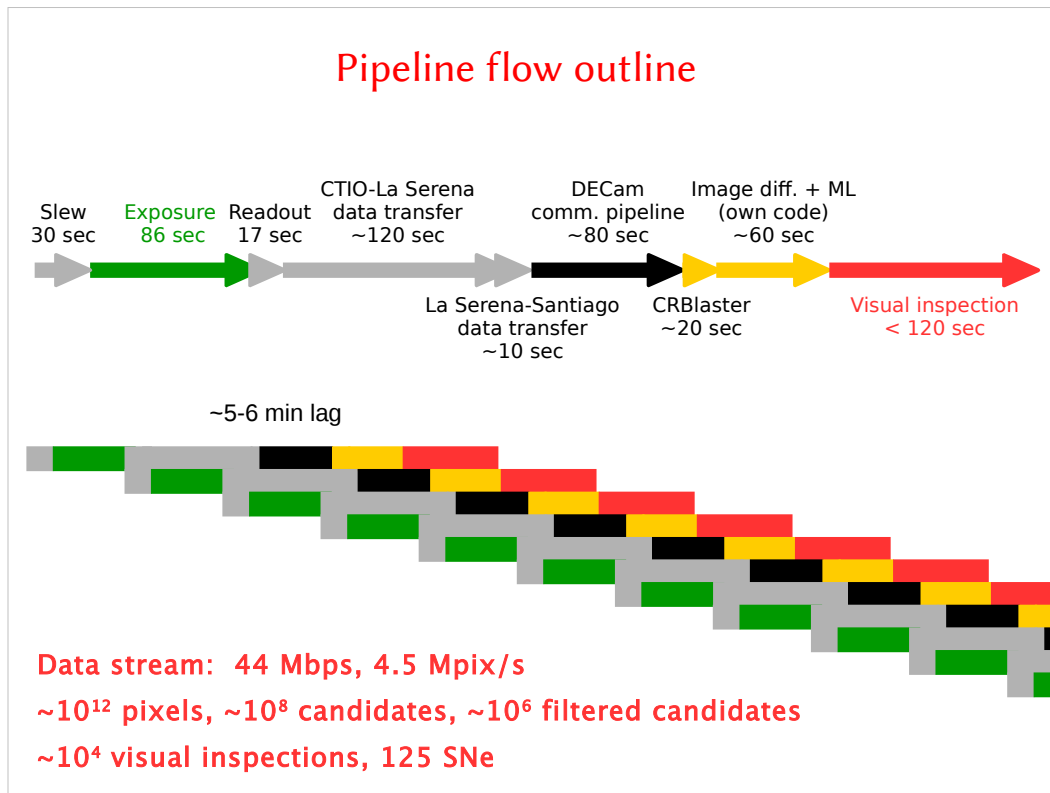
Mapa de la zona del cielo estudiada por HiTS. Cada región es observada con un pointing de DECam y contiene 520 Mpix. Cada observación toma aprox. 1.5 min. Se vuelve a la misma región 5 veces por noche.



Movimiento aparente de los campos debido a la rotación de la Tierra (de izquierda a derecha). La idea es que el telescopio permanezca en aprox. la misma posición en relación al horizonte local en cada una de las 5 épocas, i.e. que el telescopio no se mueva en cada época, sino que dejar que el cielo se mueva. Esto minimiza tiempos de movimiento del telescopio.



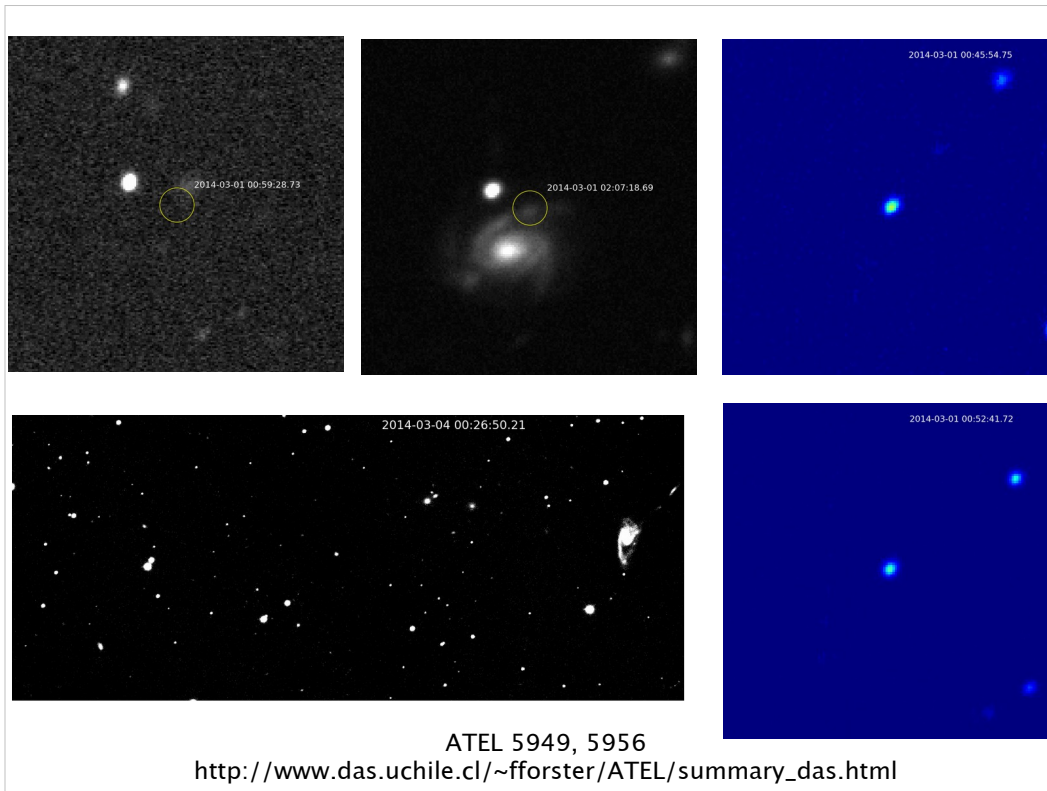
Igual que en slide anterior



Estructura de la pipeline

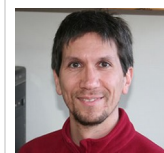
1. movimiento del telescopio (30 s)
2. tiempo de exposición (86 s)
3. tiempo de lectura (17 s)
4. transferencia por antena microondas (120 s)
5. transferencia por fibra óptica (10 s)
6. Preprocesamiento de datos en NLHPC (80 s)
7. Remoción de rayos cósmicos en NLHPC (20 s)
8. Algoritmo de alineación, convolución, diferencia, fotometría y filtrado vía señal a ruido y machine learning (nuestro algoritmo) en NLHPC (60 s)
9. Inspección visual

1e12 pix → 1e8 cand → 1e6 cand. ML → 1e4 cand
 repetidos en la misma posición (insp. Visual) → 125 SNe

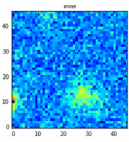


Ejemplos de arriba a la izquierda en dirección de punteros del reloj: supernova, supernova, estrella eclipsante, estrella pulsante RR Lyrae, asteroide

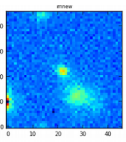
Other science



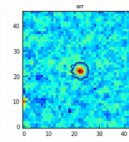
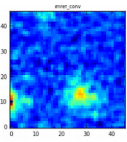
DECam



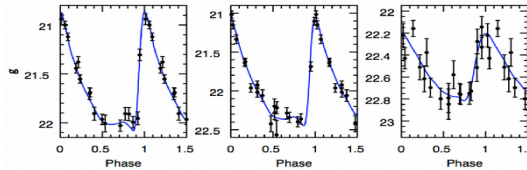
SOI



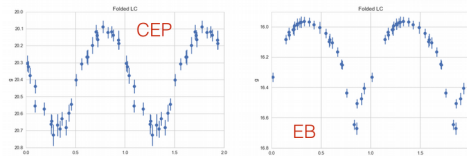
DECam conv. Difference



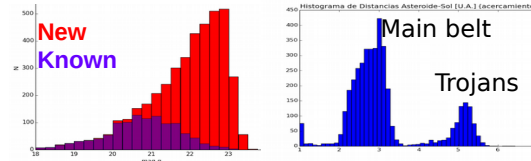
Alessandro Razza:
Combining Decam with
other telescopes



Gustavo Medina:
Most distant RR Lyrae in the
galaxy



Jorge Martínez:
Classifying every transient
object, astroinformatics



José Peña
Detecting and
characterizing asteroids

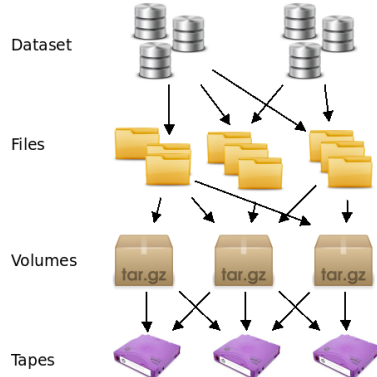
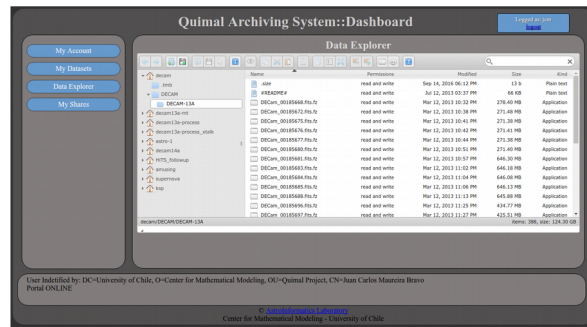
Ciencia relacionada:

1. combinación de datos DECam y otros telescopios
2. Estudio de RR Lyrae en las regiones externas de nuestra galaxia. Permite medir distancias y medir perfil de densidad de la galaxia, medir distancias a galaxias más cercanas
3. Clasificación de todas las estrellas variables usando machine learning
4. Descubrimiento y caracterización de miles de nuevos asteroides en el cinturón de asteroides y telescopios trojanos en la órbita de Júpiter

Quimal archiving system



Juan Carlos Maureira



LTO6 - 2.5 TB raw, 6.25 TB 2:5:1 compression
 LTFS - Linear Tape File system
 PKI - Public Key Infrastructure



Sistema jerárquico de almacenamiento basado en robot de cintas y discos duros con partes móviles y de estado sólido

National Laboratory for High Performance Computing (NLHPC)



Fastest computer in Chile, 2nd
in Latin America (LARTop50)
2640 cores (44 TFlops)
Infiniband (FDR 56 Gbps)
Lustre storage (DDN)

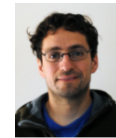
Services for:
Academia
Public sector
Industry

Team of HPC experts

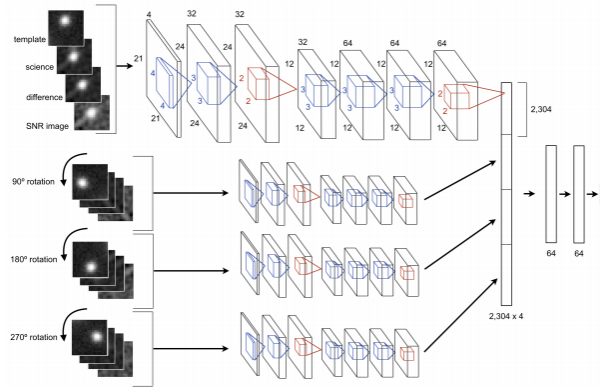
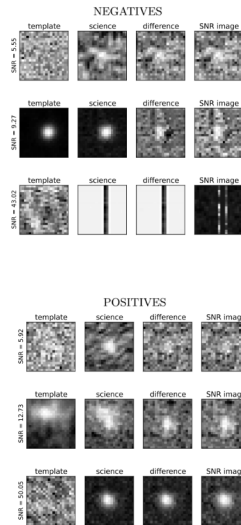
Deep learning for real bogus separation



Ignacio Reyes
DIE



Guillermo
Cabrera
CMM, U. Chile



Red neuronal convolucional profunda usada para clasificar candidatos reales vs falsos. Al entrenar esta red se definen automáticamente los features que diseñamos manualmente en una primera iteración usando random forest classifiers. La gran ventaja de esta red es que es más sensible en bajas señales a ruido, donde un humano no es capaz de desarrollar features apropiados para una buena clasificación.

Un-biased Supervised Learning



Guillermo Cabfera
CMM, U. Chile



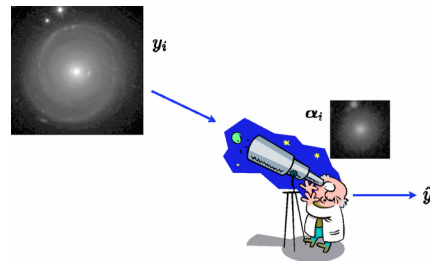
Chris Miller
Astronomy, U. Michigan



Jeff Schneider
Robotics Institute, CMU



- Biased labels due to observable parameters
- Quantify the labeling bias.
- De-bias the labels.



Estudio de bias en los set de entrenamiento y método para corregirlo.

Interferometric Image Reconstruction

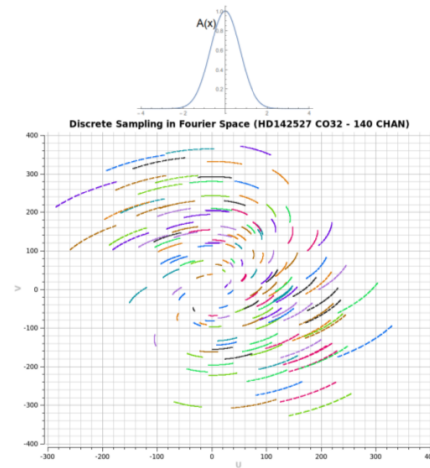


Pablo Román
Postdoc
Ph.D. Systems Eng.



Simón Casassus
Professor
Ph.D. Astrophysics

- 1 $V(u, v) = \mathcal{F}[\mathcal{A} \cdot I](u, v)$
- 2 V function is a random variable: is a Random Field.
- 3 We want: estimation for $E(I)$.
- 4 We want: estimation for $E(I^2)$.
- 5 Our data: (Simplified version)
 - Sampling points $\{z_k = (u_k, v_k)\}_{k=1}^N$,
 $N \sim 10^4 - 10^6$
 - Sampled function values at those points $\{V_k^o\}_{k=1}^N$
 - Variance estimation of measurement $\{\sigma_k\}_{k=1}^N$



Síntesis de imágenes de ALMA usando maximum entropy algorithms

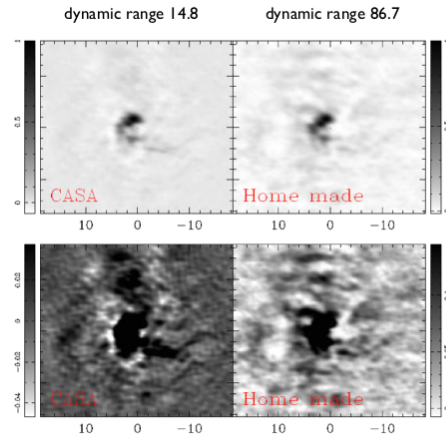
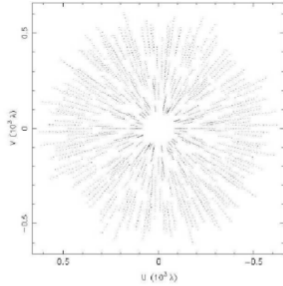
Interferometric Image Reconstruction



Pablo Román
Postdoc
Ph.D. Systems Eng.



Simón Casassus
Professor
Ph.D. Astrophysics



dynamic range: the ratio of the peak signal over the norm of the most negative artifact.

Lo mismo que slide anterior, notar mayor rango dinámico en imágenes usando máxima entropía

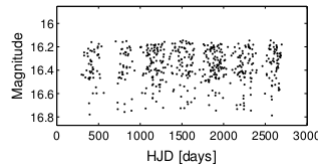
Correntropy Non-negative Matrix Factorization (NMF) Spectrum



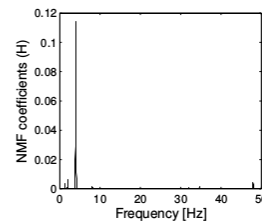
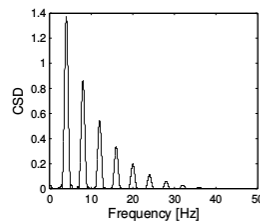
Pablo Huijse
Postdoc
Ph.D. Elec. Eng.



Pablo Estévez
Professor
Ph.D. Elec. Eng.



- Decompose correntropy into a frequency dictionary
- The NMF coefficients are the spectral ordinates of the periodogram
- Overcomplete dictionary + optimization procedure: sparse and high-res periodogram
- Adapt the kernel size of the atoms to further increase resolution



Método para detectar períodos usando non negative matrix factorization como regularizador de espectro de potencia de períodos.

Correntropy NMF Spectrum



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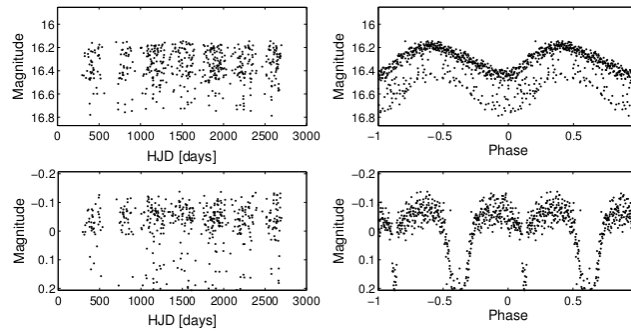
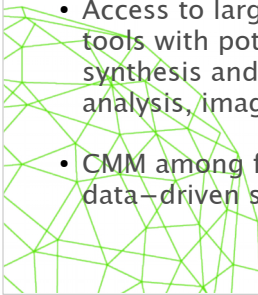


Figure: Pulsating with a 3.6 days periods plus EB with 68 days period.

Igual que slide anterior

The future

- Big opportunities ahead for Chilean astronomy, but bigger challenges for traditional astronomy → growth of astroinformatics
- Data–driven, interdisciplinary **teams** needed.
- Access to the National Laboratory for High Performance Computing a unique opportunity
- Astroinformatics must tackle real astrophysical questions and work closely with astronomers to be relevant
- Access to large volumes of data is a test bed for new data analysis tools with potential applications in other fields (e.g. ALMA image synthesis and medical imaging, LSST for real–time data stream analysis, image processing, machine learning)
- **CMM among first truly interdisciplinary centers in the world, pioneered data–driven science in Chile. Focus on technology transfer.**

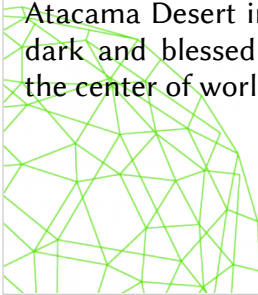


New York Times

“More Eyes on the Skies”, July 2014

“If it all plays out as expected and budgeted, astronomers of the 2020s will be swimming in petabytes of data streaming from space and the ground”

“The inauguration of these new telescopes, early in the next decade, will further enshrine the Atacama Desert in Chile, which is bone-dry, high, dark and blessed with remarkably steady air, as the center of world astronomy”



Looking Up

Northern Chile is dry and high, with little light pollution. The area already bristles with telescopes, and more are under construction.



Source: European Southern Observatory
By The New York Times

THANKS!

